

The Bamberg photographic plate archive

The digitizing project

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Point of View



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- observational astronomy
 - today: CCD detectors
 - ≥ 20 years ago: photographic emulsions (mainly on glass plates)
- plate archives
 - Sonneberg (250.000), **Bamberg (40.000)**, Hamburg (35.000), Heidelberg (15.000), Bonn (11.000), Postdam (9.000)
- Bamberg Southern Photographic Patrol Survey (Strohmeier 1963)
 - Boyden, South Africa, 1963-1976, aim: search for variable stars
 - direct images, $B \leq 14.5$ mag, $16\text{cm} \times 16\text{cm} = 14^\circ \times 14^\circ$
 - **all southern data from this time is primarily based on this survey**

Need for Preservation



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■ historian's point of view:

- oldest plates at Bamberg from 1910s
- photographic emulsions already begin to decompose and dissolve from their glass plates

■ scientist's point of view:

- former results using blink comparators:
→ 1.700 'Bamberg variables'
(e.g. Strohmeier & Knigge 1961, Friedrich & Schöffel 1971,
Knigge & Sosna 1977)
- only a fraction of observed plates have been looked on
- 25.000 - 200.000 stars on one plate
→ just 'tip of the iceberg'

Need for Preservation



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- equipment became affordable
 - large amount of storage space
- give access to the community (Virtual Observatory)

⇒ **digitizing project** (DFG funded)
in collaboration with the Leibniz Institute for Astrophysics at
Potsdam (AIP) and the Hamburg Observatory

Metadata



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■ 11 years ago:

start of project to incorporate all
metadata into Wide-Field Plate
Database (WFPDB)

(Tsvetkov et al. 2005,
Tsvetkova et al. 2006)

→ www.skyarchive.org

WFPDB search page

1

**Wide-Field Plate Database - Sofia
Search Page**

WFPDB Archives Active WFPDB Archives

Query Setup

Maximum entries per page: 50 Output layout: htm table Output Order: + - Reset A

Search by Object or Field Coordinates

RA J2000 hh mm ss Star Gazer Visualization
DEC J2000 dd mm ss
Field Size deg
 Instrument Field Radius Box Size
 Reduced Instrument Field

More constraints

Plates with previews only Magnitude Limit _____

Additional display

Angular Distance from Field Centre Preview (if available) JD

Select A

Search by Constraints applied on Columns

Show	Sort	Column	Constraints	Units	Explain
<input checked="" type="checkbox"/>	<input type="radio"/>	IDobs	bam	(char)	WFPDB observatory identifier
<input checked="" type="checkbox"/>	<input type="radio"/>	IDins	010	cm	Instrument aperture
<input checked="" type="checkbox"/>	<input type="radio"/>	-	DSuf1	(char)	Instrument aperture suffix
<input checked="" type="checkbox"/>	<input type="radio"/>	-	IDno	(char)	Original plate number
<input checked="" type="checkbox"/>	<input type="radio"/>	DSuf2		(char)	Original plate number suffix
<input checked="" type="checkbox"/>	<input type="radio"/>	RAJ2000		hh:mm	Right ascension
<input checked="" type="checkbox"/>	<input type="radio"/>	DECJ2000		dd mm	Declination
<input checked="" type="checkbox"/>	<input type="radio"/>	-	CCOD		Code for Error, Missing Data, or Uncertainty of coordinates
<input checked="" type="checkbox"/>	<input type="radio"/>	DATE/UT		YYYY MM DD HH:MM:SS	Date/time of observation
<input checked="" type="checkbox"/>	<input type="radio"/>	-	TCOD		Code for Error, Missing Data, or Uncertainty of observation time
<input checked="" type="checkbox"/>	<input type="radio"/>	-	OBJNAM	(char)	Object or field designation
<input checked="" type="checkbox"/>	<input type="radio"/>	-	OBJTYP	(char)	Object type code
<input checked="" type="checkbox"/>	<input type="radio"/>	-	METHOD		Method of observation code
<input checked="" type="checkbox"/>	<input type="radio"/>	-	MULTEX		Multiplicity of exposure
<input checked="" type="checkbox"/>	<input type="radio"/>	-	EXP	min	Exposure time
<input checked="" type="checkbox"/>	<input type="radio"/>	-	EMULS	(char)	Emulsion type
<input checked="" type="checkbox"/>	<input type="radio"/>	-	FILT	(char)	Filter type
<input checked="" type="checkbox"/>	<input type="radio"/>	-	SPEC	(char)	Spectral band
<input checked="" type="checkbox"/>	<input type="radio"/>	-	DIMx		X dimension of plate
<input checked="" type="checkbox"/>	<input type="radio"/>	-	DIMy		Y dimension of plate
<input checked="" type="checkbox"/>	<input type="radio"/>	-	PQUAL		Link to quality information
<input checked="" type="checkbox"/>	<input type="radio"/>	-	PNOT		Link to note text
<input checked="" type="checkbox"/>	<input type="radio"/>	-	POBS		Link to observer's name(s)

<http://wfpdb.org/search/>

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Scan



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- flatbed scanner
EPSON Expression
10000XL
 - resolution:
 $2.400 \text{ dpi} = 10 \mu\text{m} =$
 $3''.6/\text{pix}$
 - duration: 5 minutes
 - size: 400 MByte



Data Preparation



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- implementation in database (Virtual Observatory)
 - all scanned data (inclusive logbooks) + metadata
 - FITS-header (Wide-Field Plate Database)
 - storage, web based interface
 - Leibniz Institute for Astrophysics at Potsdam (AIP)
- scientific investigations
 - ⇒ **data extraction**
 - astrometry
 - photometry

Astrometry

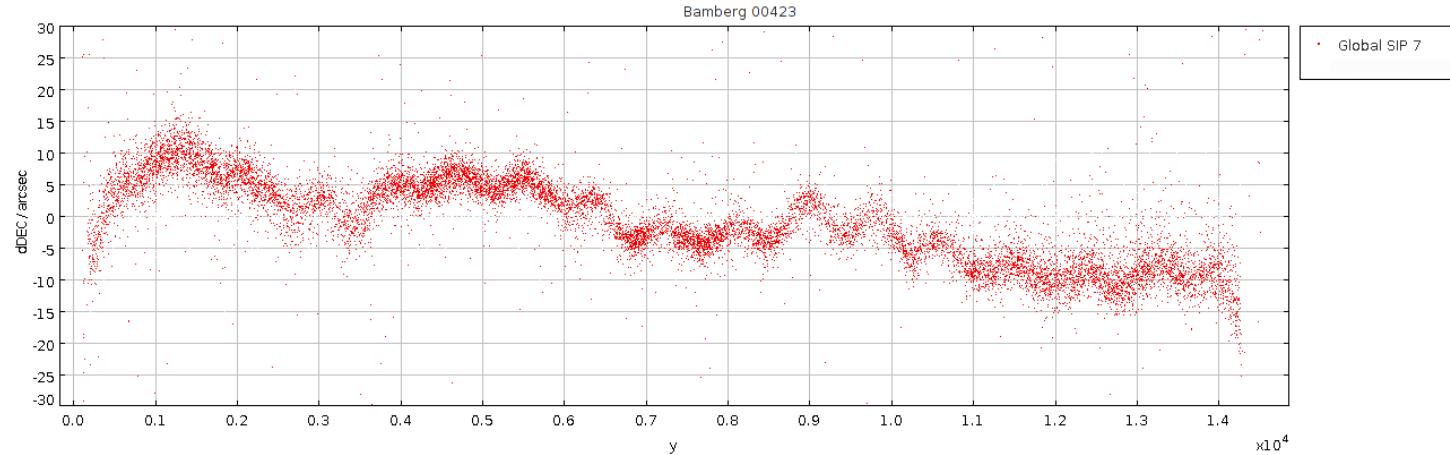


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- SExtractor $\longrightarrow x, y$
- astrometry.net \longrightarrow RA, DEC



- ⊕ no additional file

FITS header: Simple Imaging Polynomial (SIP)

- ⊕ $\langle \text{RA}, \text{DEC} \rangle = \pm 15''$
- ⊕ for good solution \rightarrow **2** scans necessary

Astrometry

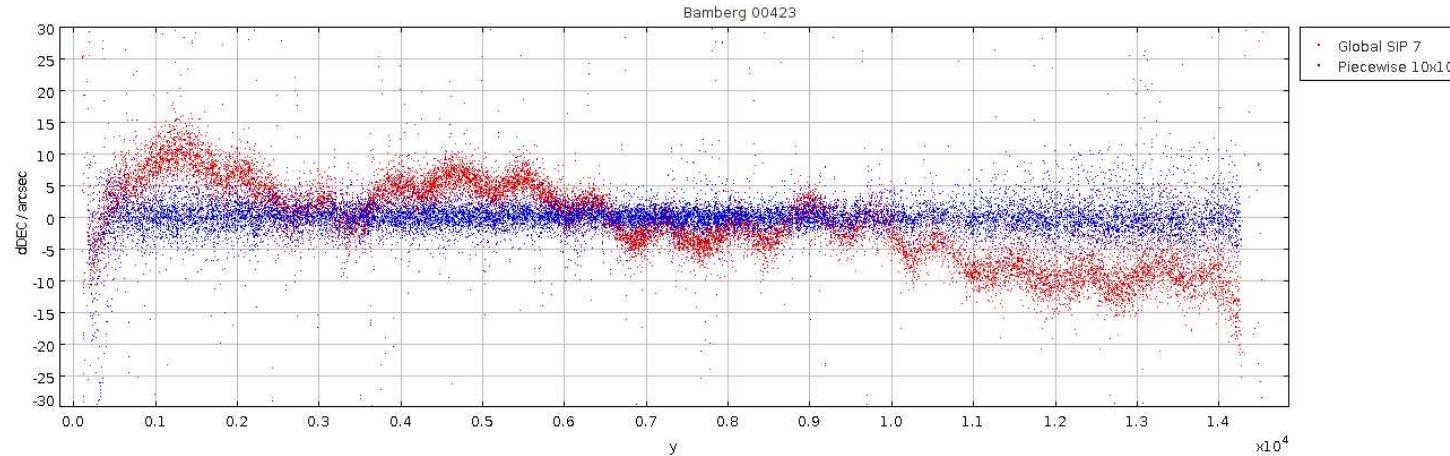


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- SExtractor $\rightarrow x, y$
- astrometry.net, SCAMP + segmentation \rightarrow RA, DEC



⊕ $\langle \text{RA}, \text{DEC} \rangle = \pm 2''$ (note: scan resolution = $3''.6/\text{pix}$)

→ less than expected → search for e.g. asteroids, comets

⊕ for good solution \rightarrow 1 scan necessary

⊖ additional file (source list)

Photometry



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- SExtractor → relative magnitudes (using Zero point value)
- calibration: catalogs → magnitudes

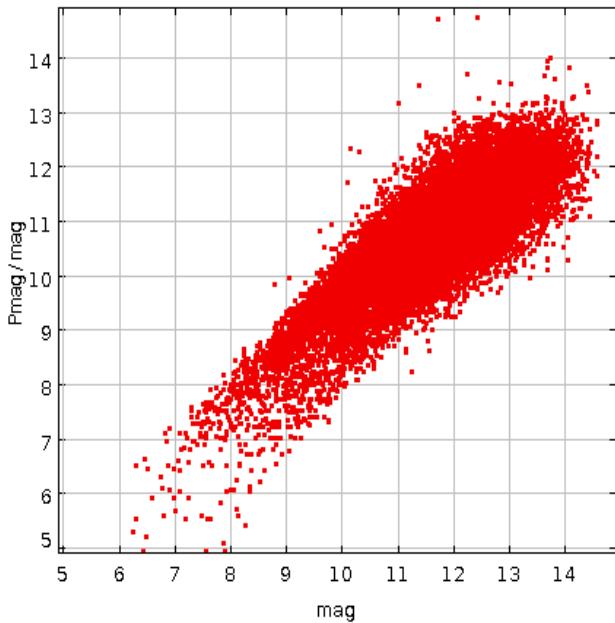


plate vs. HST GSC

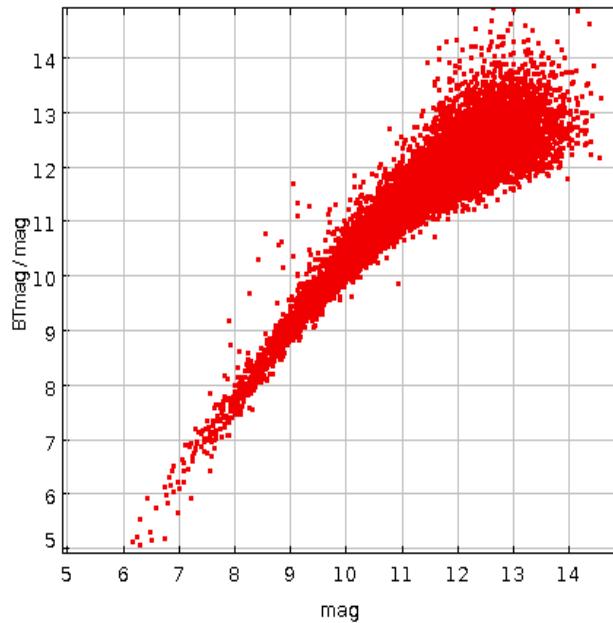


plate vs. Tycho-2

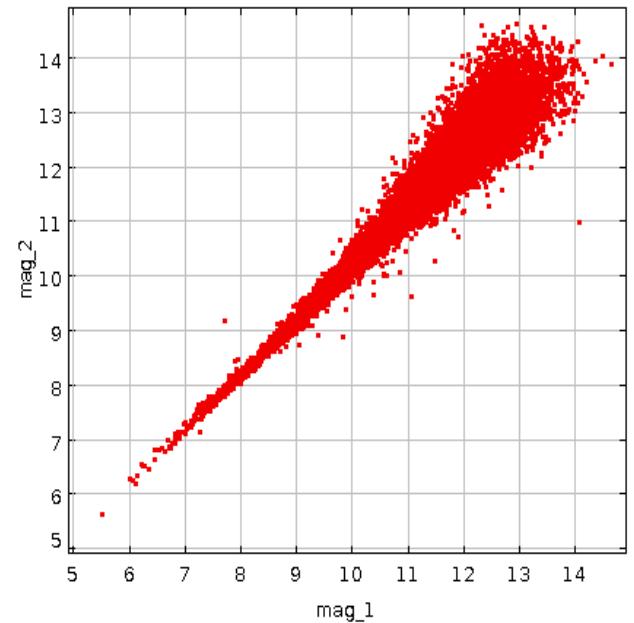


plate1 vs. plate2

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- SExtractor → relative magnitudes (using Zero point value)
- calibration: catalogs → magnitudes

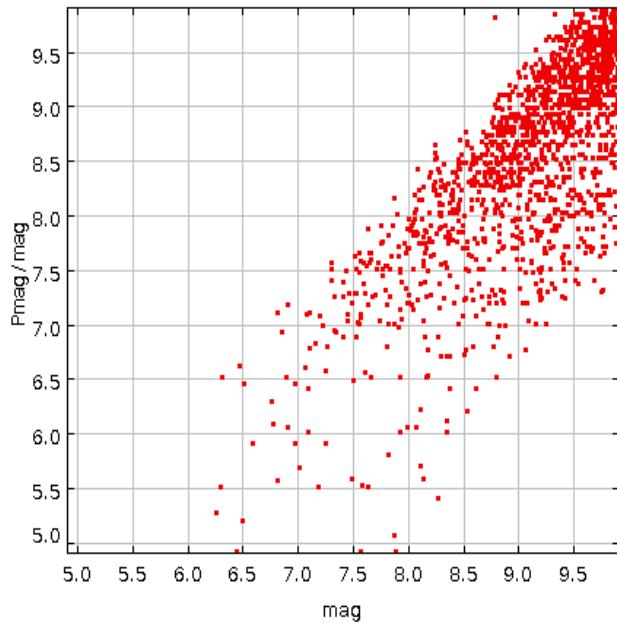


plate vs. HST GSC

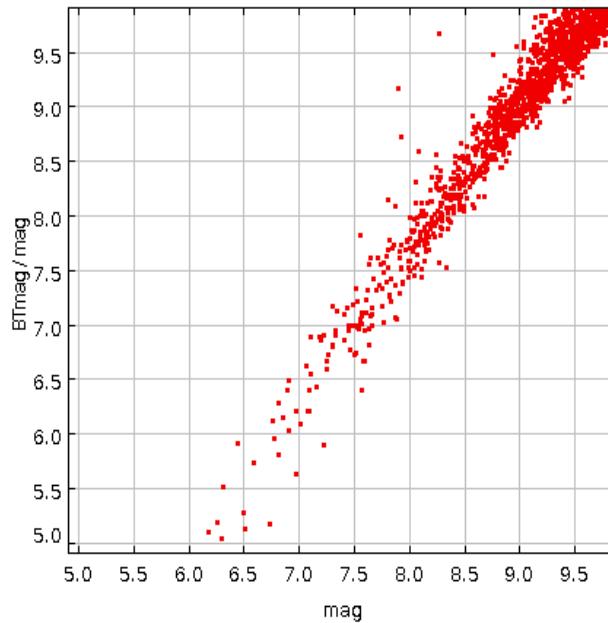


plate vs. Tycho-2

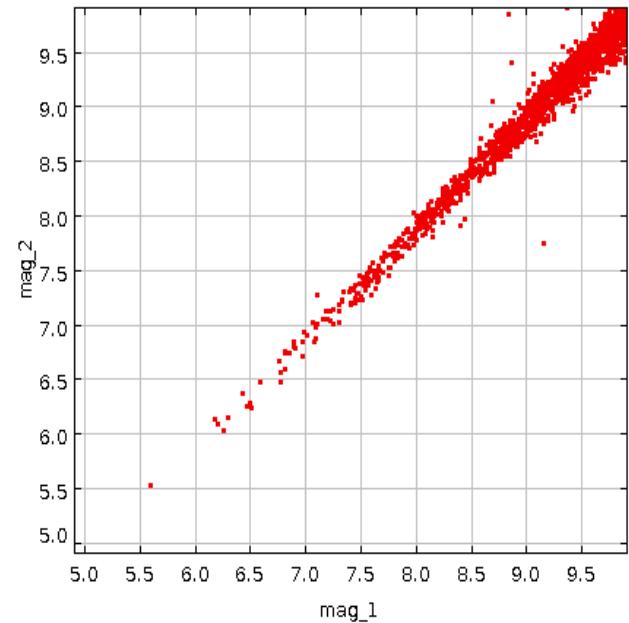


plate1 vs. plate2

First Tests

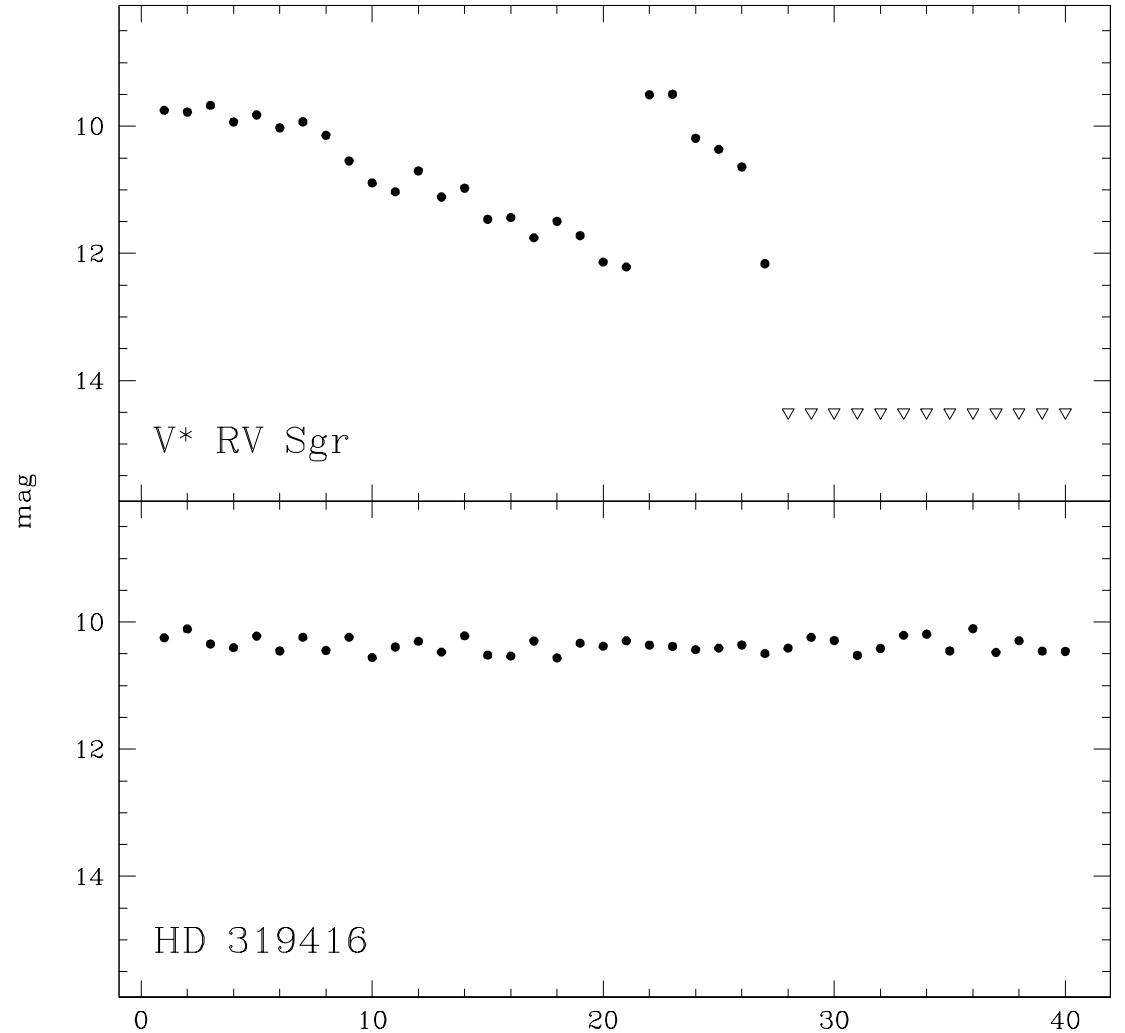


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- known Mira
- $8\text{mag} \leq V \leq 13\text{mag}$

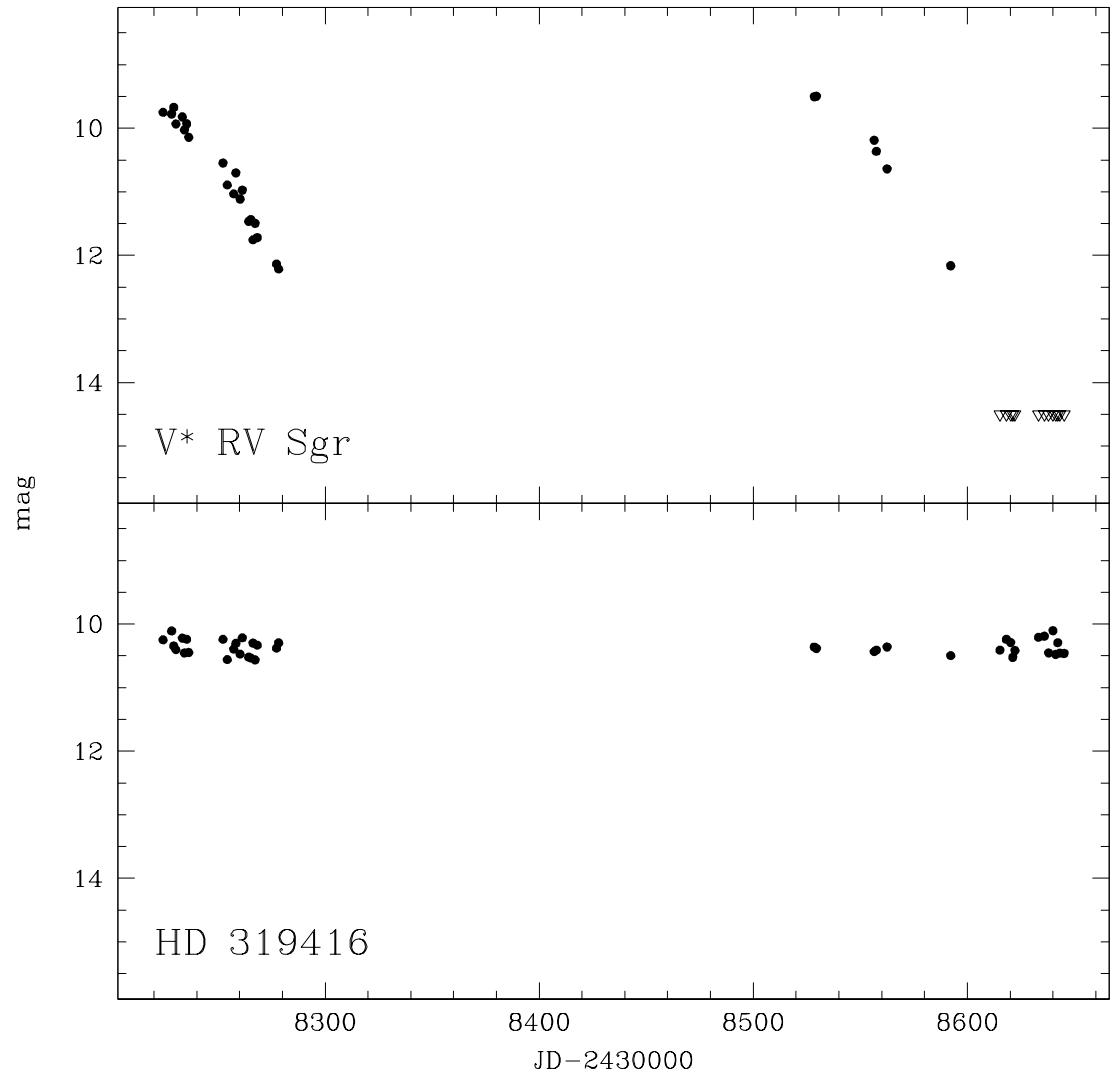


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- known Mira
- $8\text{mag} \leq V \leq 13\text{mag}$



References



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