

Development of an Open Source light curve classificator

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University part

- Master student (physics) @University of Marburg
- Bachelor in 2013: "Photometry at Wolf-Rayet-Stars"
- Currently working on master thesis "Knowledge Discovery in the OGLE-III database of variable stars"

Community part

- Developer @Fedora Linux
- ▶ Working in groups for Astronomy, Science & Technology
- Including packages like astroML or Astromatic Tools in Fedora



Astronomy Group at University of Marburg



Gerling Observatory

Research group

- History of Astronomy and Observational Astronomy
- Prof. Dr. Andreas Schrimpf

Research topics

- Analysis of Sonneberg Plate Archive
 - Poster "First Steps Towards a Photometric Analysis of the Sonneberg Sky Patrol Plates" by M. Spasovic
- Data Mining in context of variable stars
- History of Astronomy (Christian Ludwig Gerling)

In cooperation with

- Sonneberg Observatory (P. Kroll)
- Department of Physics & Astrophysics, University of Delhi, India (H. P. Singh)



About this Talk

Concept:

- Presentation of a concept for development of a classificator
- Based on experiences from work @master thesis
- \blacktriangleright Work in progress \Rightarrow Focussed on preparation of data

What we want to have

- Reliable classification of light curves
- Different methods depending on ambition
- Here: Machine Learning (ML)
 - Supervised: Filter for known classes
 - Unsupervised: Search for new (sub-) classes
- Reproducible results



State of the Art - Light curve analysis



From measurement to final results



- Look at (c): "Oh, looks like an anomalous Cepheid???"
- Output of step (c) \Rightarrow Machine Learning
- Interpret output of ML-based analysis



The input data



- Photometrical data reductions done
- Problems
 - Data unequeally spaced at time axis
 - \Rightarrow Not a direct input for ML
- We have to calculate classifiable data from input



How to solve?



Assumption (for master thesis): periodic variable stars

► Idea: Fold all measured data in one "period" and analyze mag(φ(t)) instead of mag(t)

$$\phi\left(t\right) = \frac{t - t_0}{P} - \operatorname{Int}\left(\frac{t - t_0}{P}\right)$$

- Result: We loose dependency on date of measurement and get characteristic light curves
- Standard technique in field of variable stars



Generating classifiable data

Problem: Data still unequally spaced and not normalized

- Solutions:
 - Normalize mag into interval [0,1]
 - Fit a polynomial to the normalized phased light curve
 - Use fit parameters for classification
 - Generate a synthetic equispaced light curve using fitted function
 - Fit Fourier parameters R_{ij}, ψ_{ij} to light curve and use them for classification

$$m(t) = A_0 + \sum_{n=1}^{N} A_n \cos (2\pi n\phi (t) + \psi_n)$$
$$R_{ij} = \frac{A_i}{A_j} \qquad \psi_{ij} = \psi_i - i\psi_j$$

Fourier-based classification (without ML) established method



Importance of period search implementation



- Slightly different period: 3.5 seconds, P about 20.4 hours
- Calculated period with FNPEAKS¹
- Calculated Fourier parameters using OGLE and FNPEAKS periods:

	R ₂₁	R ₃₁
OGLE	0.460 ± 0.014	0.364 ± 0.013
FNPEAKS	0.379 ± 0.039	0.219 ± 0.038

¹http://helas.astro.uni.wroc.pl/deliverables.php?lang=en&active=fnpeaks



Measure quality

- Difference in Fourier parameters show: We need some kind of quality measurement
- Idea: Compare standard deviation of phased light curve with error of measurements
- Classification of generated equispaced light curve more reliable

After data preprocessing and optional quality measurement: Ready for classification!



Classification



Figure: Different classifiers implemented by scikit-learn²

- > Data now viable for analysis by common ML software
- ML toolboxes typically provide implementations for
 - Supervised Learning
 - Unsupervised Learning
 - Principal Component Analysis (PCA)
 - Clustering

²http://scikit-learn.org/stable/index.html







- Optical Gravitational Lensing Experiment
- Photometrical measurements
- OGLE-III catalog of variable stars (approx 400 000 stars)
- Provides "ready to use" light curves
- Well analyzed, also ML-based

Nice base for development of a light curve classificator



ML-based classification of OGLE-data in LMC/SMC



Deb, Singh (2009): "Light curve analysis of Variable stars using Fourier coefficients and Principal Component Analysis"



Implementation



Develop as free software

- Better reproducibility by third parties
- Reasonalbe algorithms
- Available for whole community
- Many basic packages (Python + modules) are free software too
- Avoids the "reinventation of the wheel"



How to implement? Use Python!

Which language shall we use?

- ► In Astronomy community: Python is a common programming language
- astropy, numpy, scipy, …
- $\blacktriangleright \Rightarrow$ New software should at least have an interface to Python
- ▶ Guru: C/C++ is much faster...

Already done (just to be rewritten in a distributable way ...)

Data preprocessing using numpy/scipy

Available

- scikit-learn (machine learning package for Python): PCA, Clustering etc. implemented
- Python wrappers for shogun (fast C++ machine learning toolbox)

Solution: Create modular software, write critical components using C++ and generate wrappers for them (swig)



astroML



- Machine Learning and Data Mining for Astronomy
- Many tasks already implemented (periodogram for example)
- Based on
 - Python
 - scikit-learn, numpy, matplotlib
- Astropy Affiliated Package
- astroML-addons: example for integration of critical components implemented in C++

Goal: Implement light curve analysis in a way, that we can contribute it to astroML instead of initiating a new project



Conclusion

- Implement data preprocessing (draft already done)
- Quality measurement
- Test classification using scikit-learn
- Use OGLE as test data, compare with previous analysis
- Document well and publish as free Software



Thank you for your attention!

Any questions?



References

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