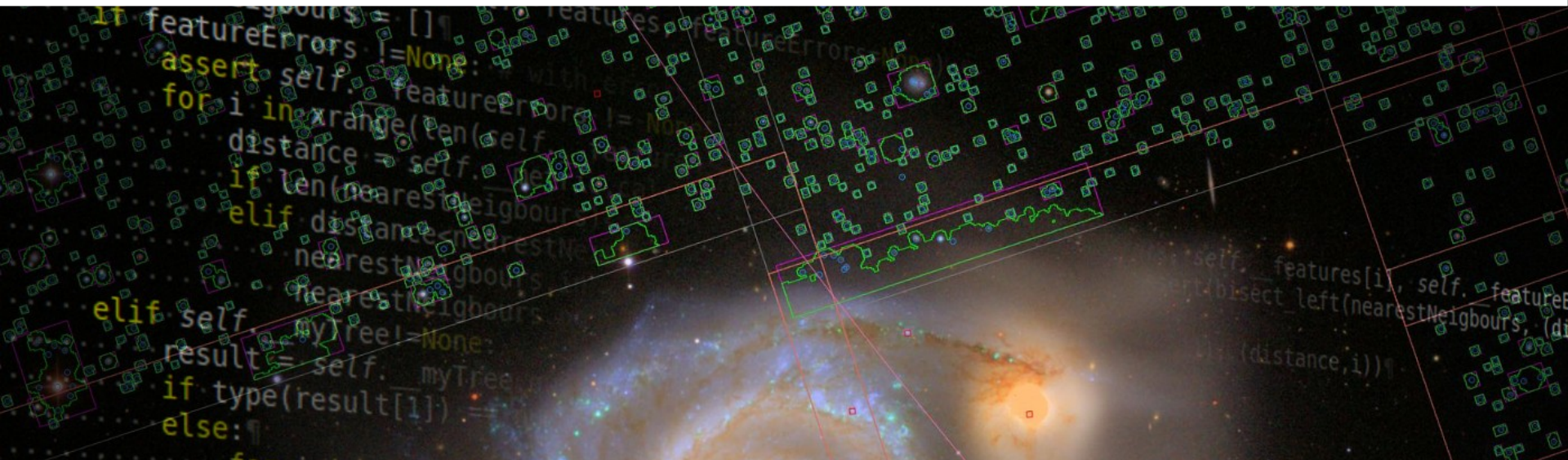
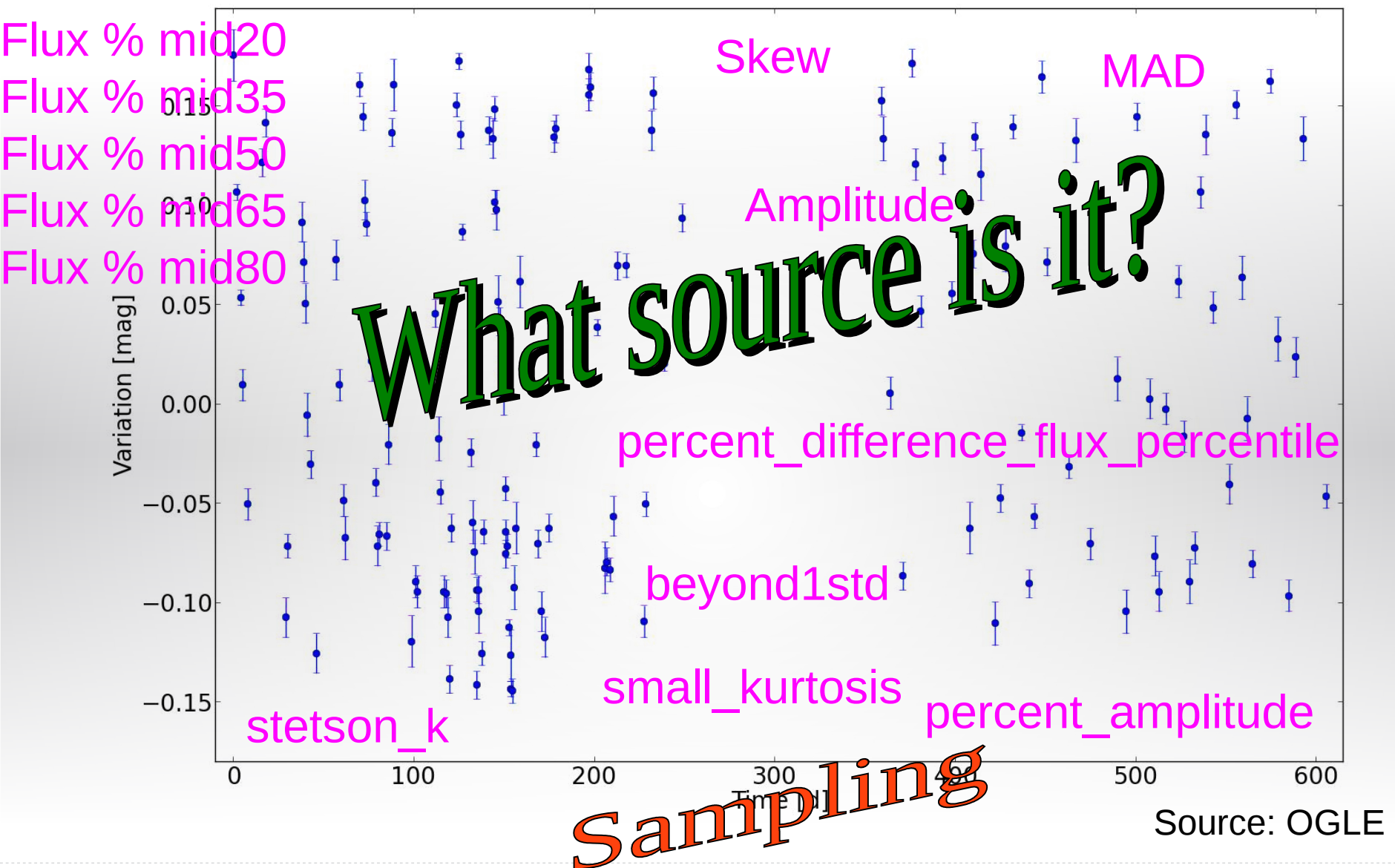




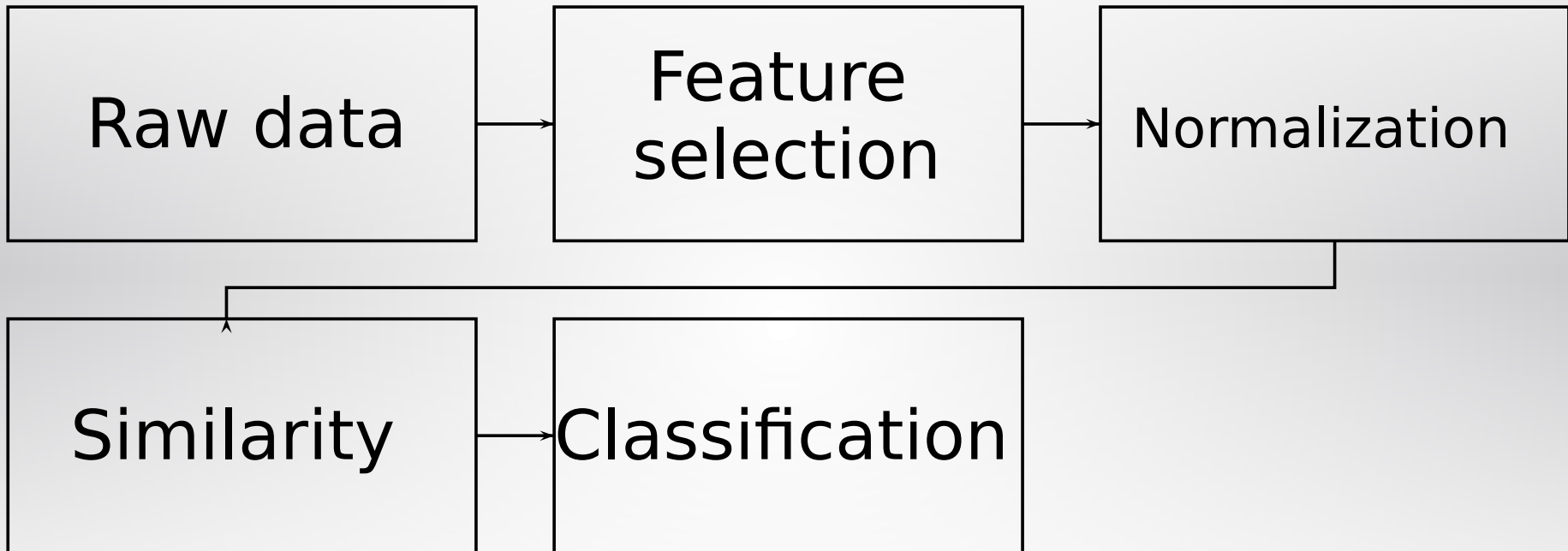
Featureless Classification of Light Curves

Sven Dennis Kügler (Astroinformatics@HITS)

Feature-based representation



Feature-based representation



Feature-based representation

Problem:

- Creation & selection of features is **arbitrary**
- Features are strongly **biased**
- Features do not allow meaningful use of **unsupervised** tasks
- **Uncertainty** cannot be taken into account



works well for classification of similarly bright objects in individual surveys

But nothing else

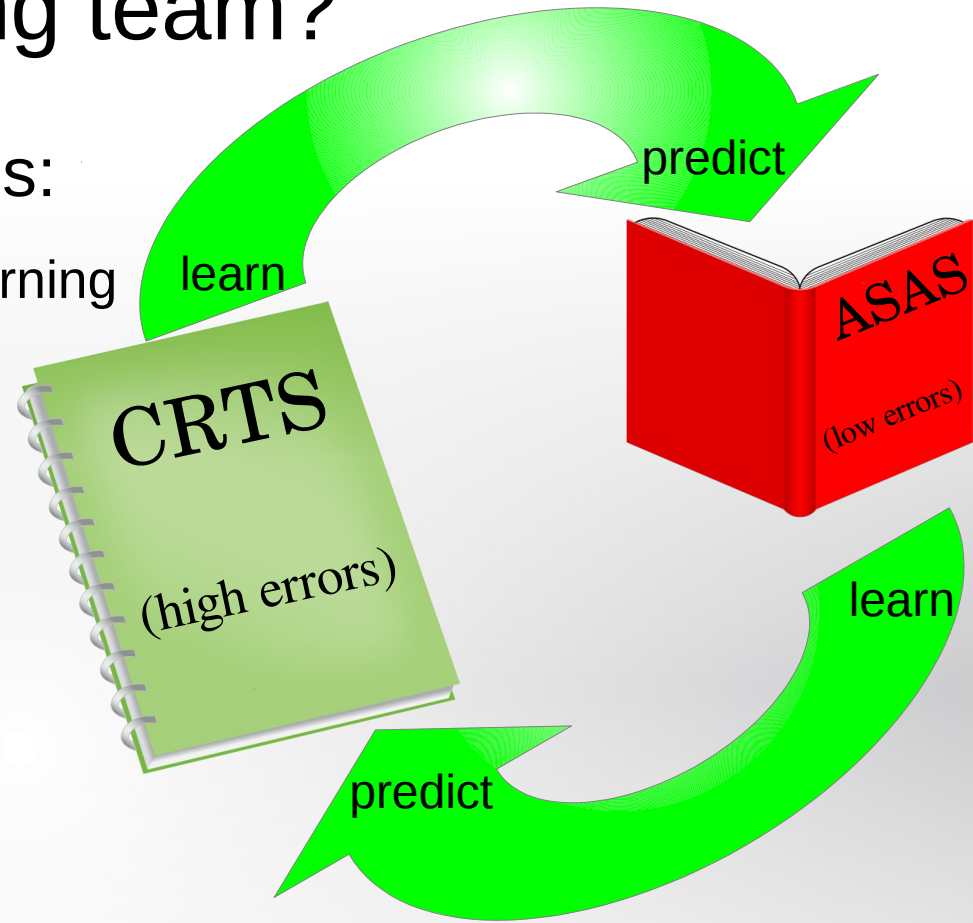
Never change a winning team?

Prospects for new approaches:

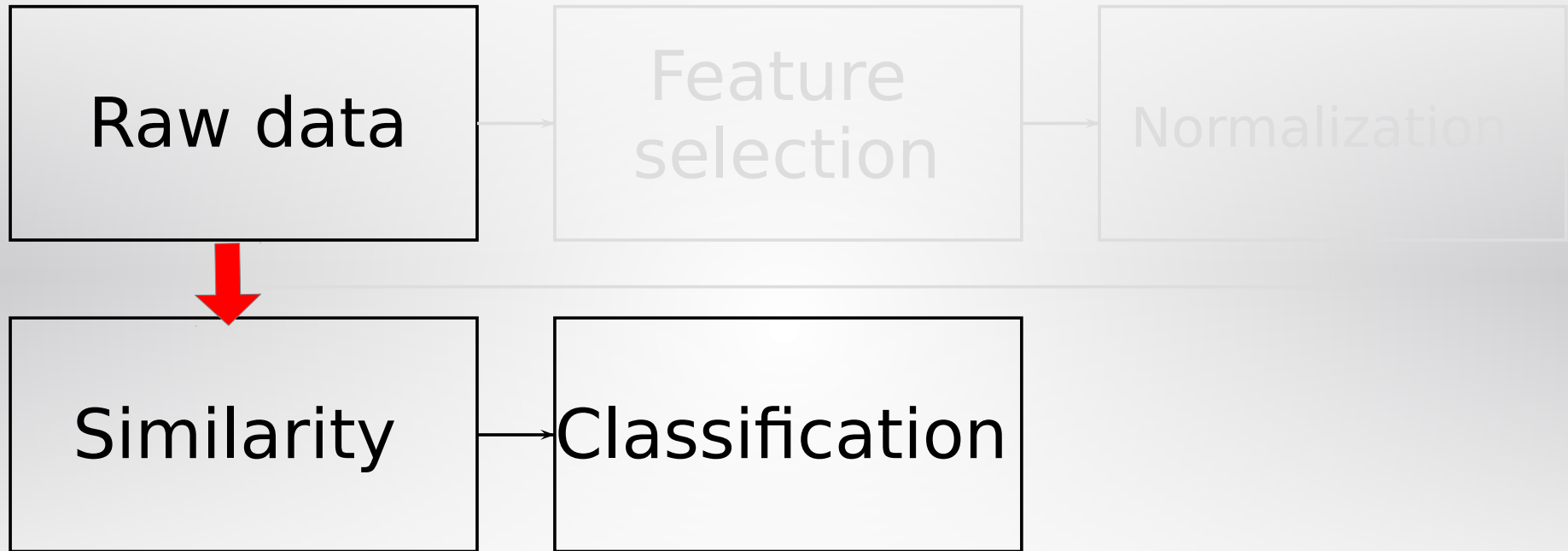
- (semi-supervised) transfer learning
- Avoiding in-survey biases
- Detection of outliers
- Unsupervised methodology



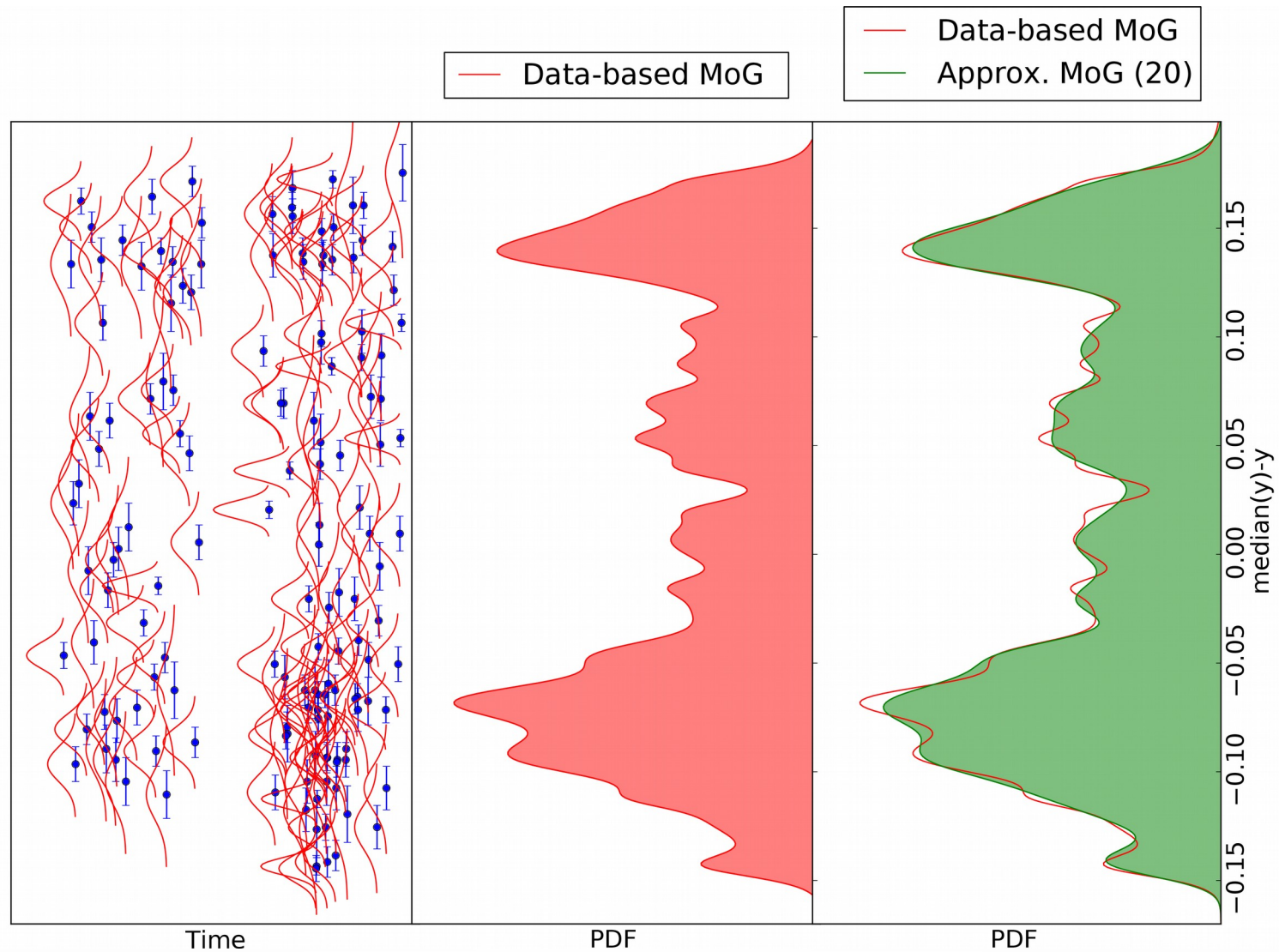
Avoiding black-box-like selections allows to obtain “real” understanding of data



Density-based representation



Density-based representation



Kügler et al., 2015, MNRAS

Density-based representation

Judging distances between two densities:

L2 norm

$$\int_{-\infty}^{+\infty} (|p(x) - q(x)|)^2 dx$$

Bhattacharyya

$$-\ln\left(\int_{-\infty}^{+\infty} \sqrt{p(x)q(x)} dx\right)$$

Kullback-Leibler-Divergence

$$\int_{-\infty}^{+\infty} p(x) \log\left(\frac{p(x)}{q(x)}\right) dx + \int_{-\infty}^{+\infty} q(x) \log\left(\frac{q(x)}{p(x)}\right) dx$$

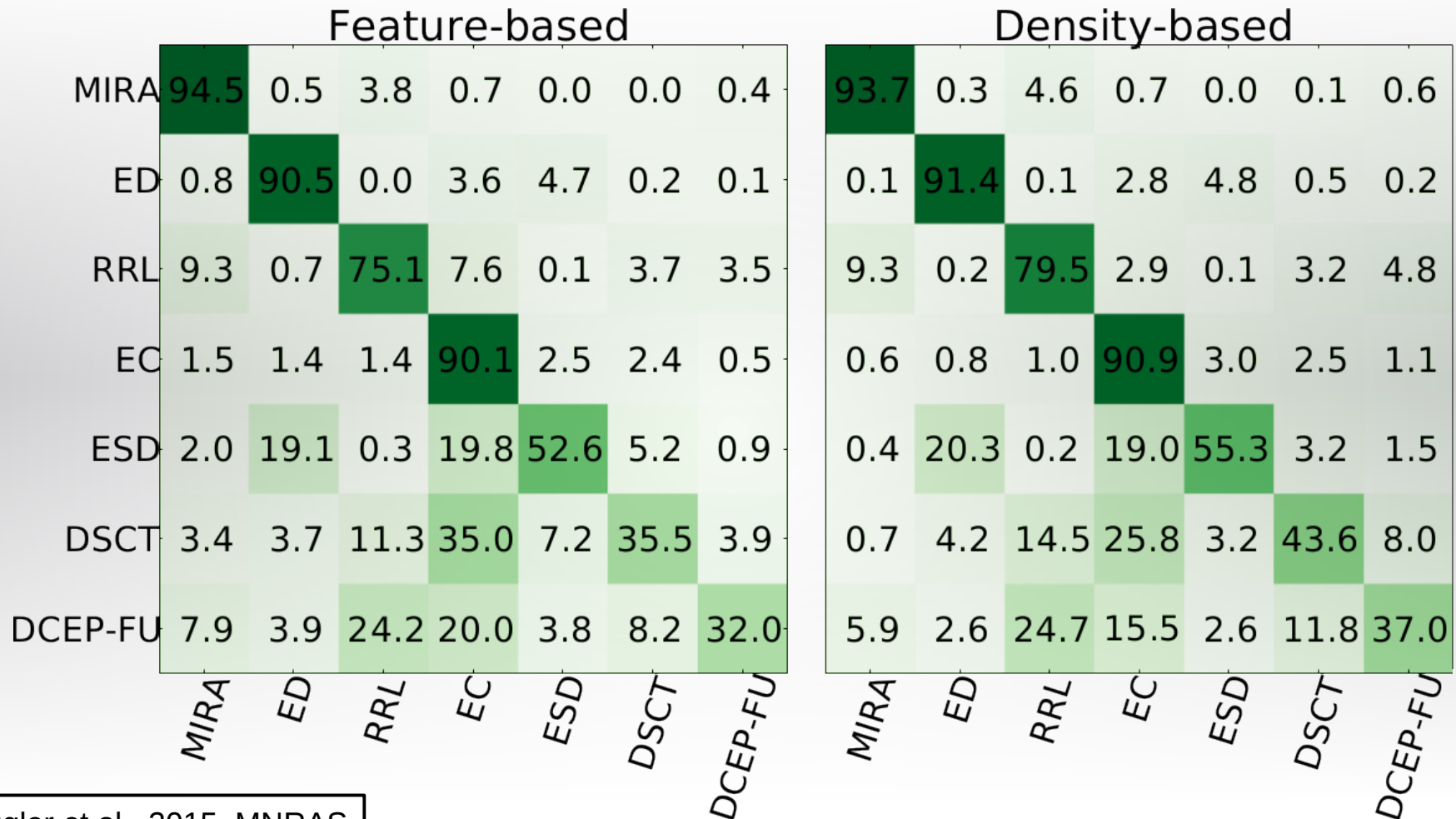
Numerical experiments

7-class-classification experiment (ASAS)

	kNN	v-SVM	RF
Features (raw)	74.22 $\pm$ 1.24 k=11	78.02 $\pm$ 0.68 ν = 0.19, δ = 0.53	79.98 $\pm$ 1.16 T=400
Features (norm.)	77.60 $\pm$ 0.76 k=17	80.47 $\pm$ 1.21 ν = 0.17, δ = 0.10	79.99 $\pm$ 1.55 T=400
L2	79.57 $\pm$ 0.80 k=19	82.08 $\pm$ 0.89 ν = 0.01, δ = 0.56	— —
KLD	78.96 $\pm$ 1.87 k=23	75.56 $\pm$ 0.94 ν = 0.26, δ = 0.34	— —
BHA	79.73 $\pm$ 0.83 k=29	81.11 $\pm$ 0.90 ν = 0.20, δ = 0.14	— —

Numerical experiments

7-class-classification experiment (ASAS)



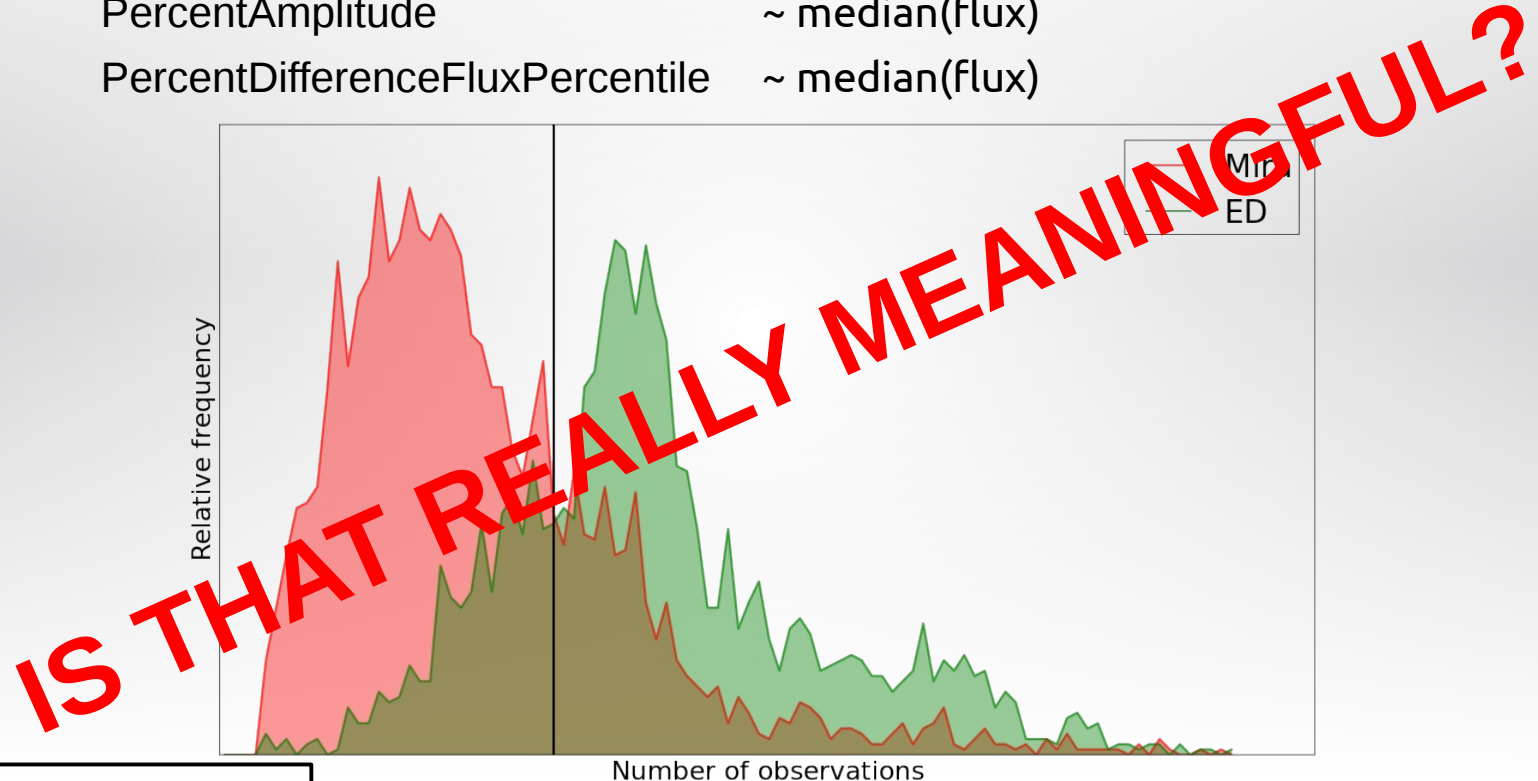
Kügler et al., 2015, MNRAS

Numerical experiments

7-class-classification experiment (ASAS)

3 features cannot be encoded in density-representation:

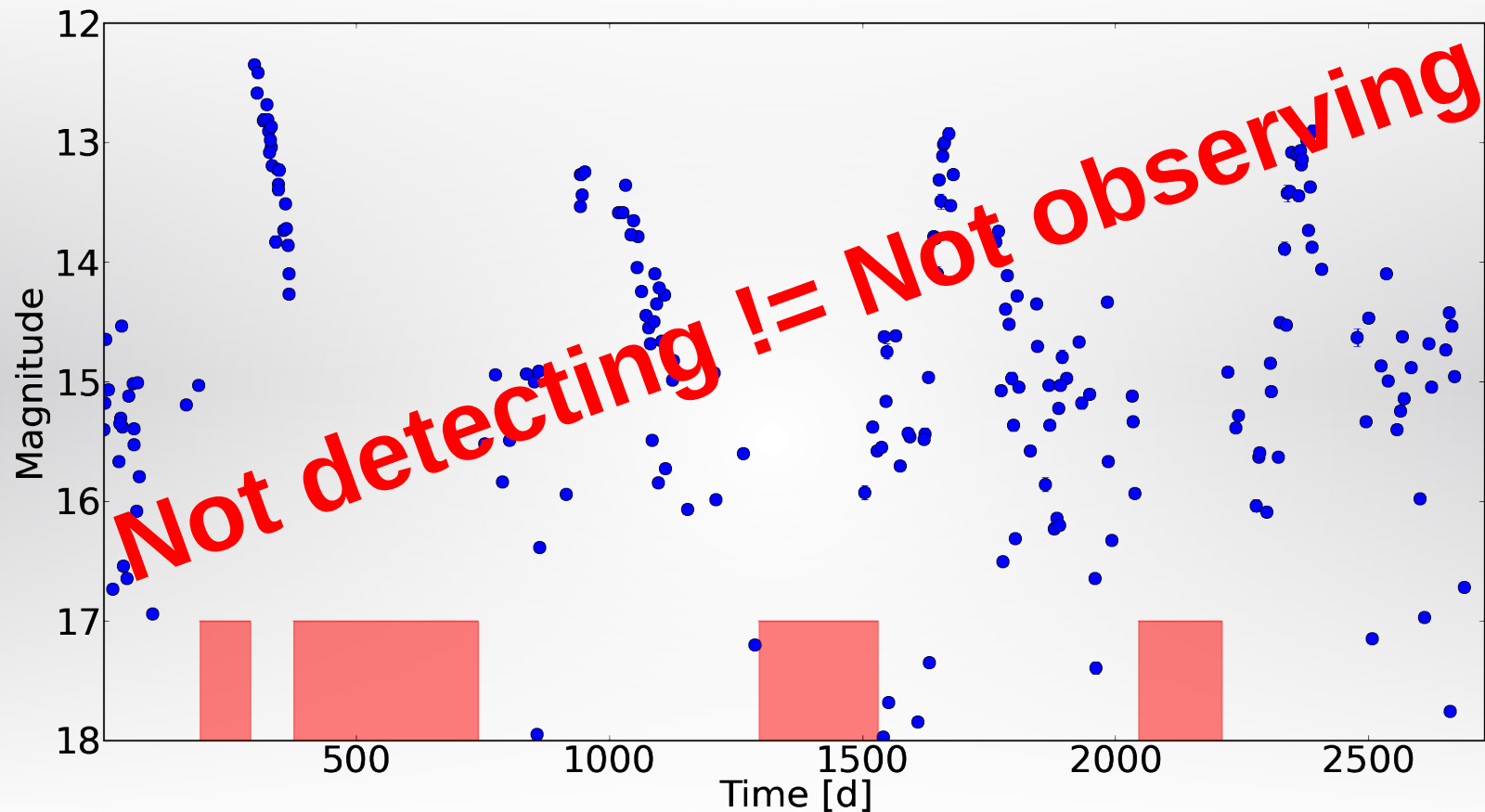
StetsonK	$\sim 1/(\text{number of observations})$
PercentAmplitude	$\sim \text{median}(\text{flux})$
PercentDifferenceFluxPercentile	$\sim \text{median}(\text{flux})$



Kügler et al., 2015, MNRAS

Numerical experiments

7-class-classification experiment (ASAS)



Analysis of Regular Time Series

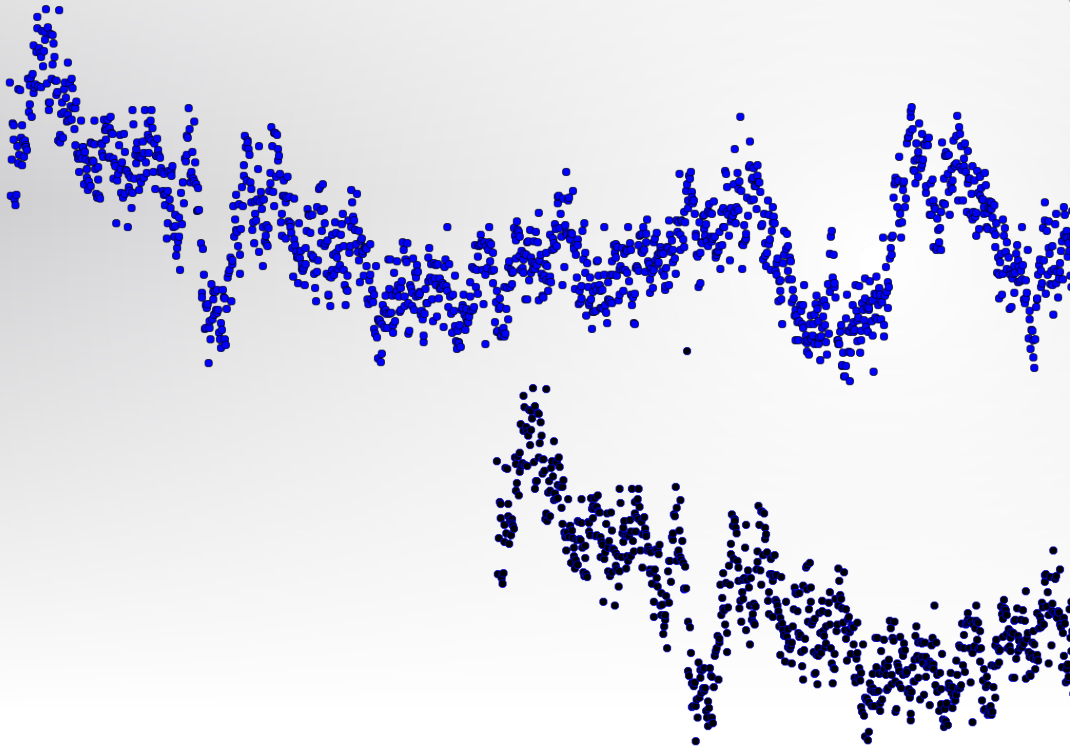
Problem:

- Time series generally of variable length
- Respect of sequential behavior
- Invariance against time shifts required

Conversion to a

vector

representation



ESN



$$\begin{bmatrix} a_1 \\ a_2 \\ \cdot \\ \cdot \\ \cdot \\ a_N \end{bmatrix}$$

ESN



$$\begin{bmatrix} b_1 \\ b_2 \\ \cdot \\ \cdot \\ \cdot \\ b_N \end{bmatrix}$$

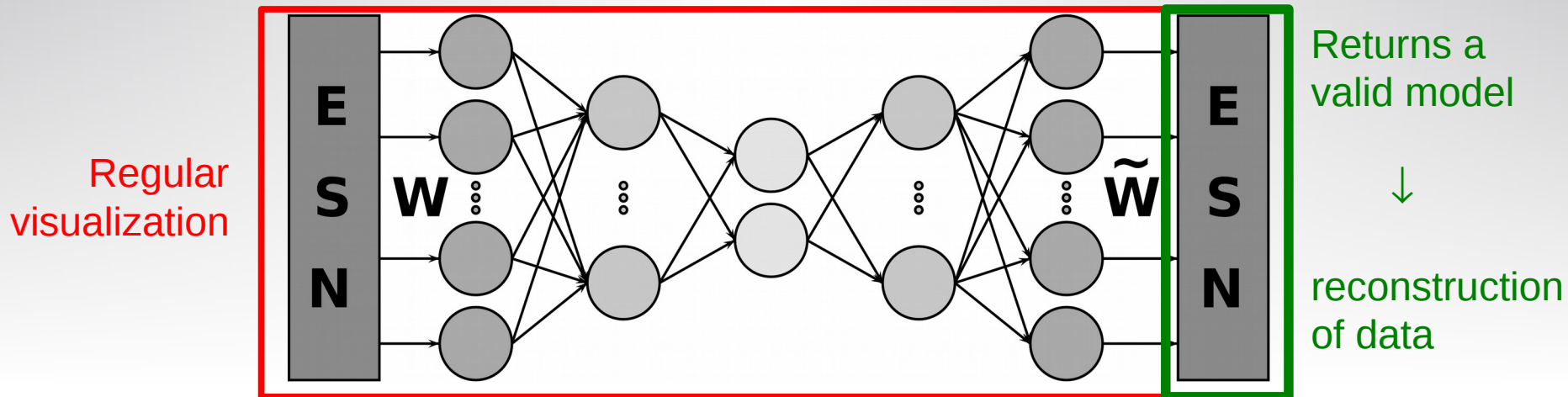
Analysis of Regular Time Series

These **vector representations** can be visualized.

However

- Visualization algorithms: optimal de- & compression of representation
- Scientifically more interesting: reconstruction error on the data

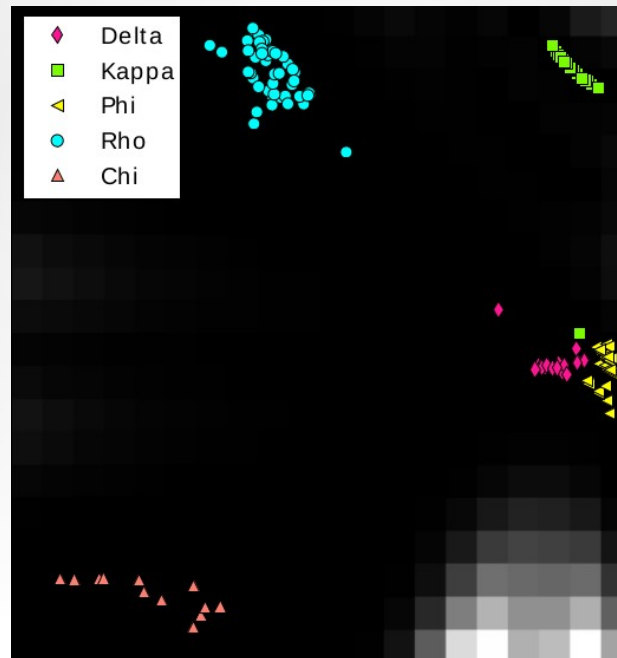
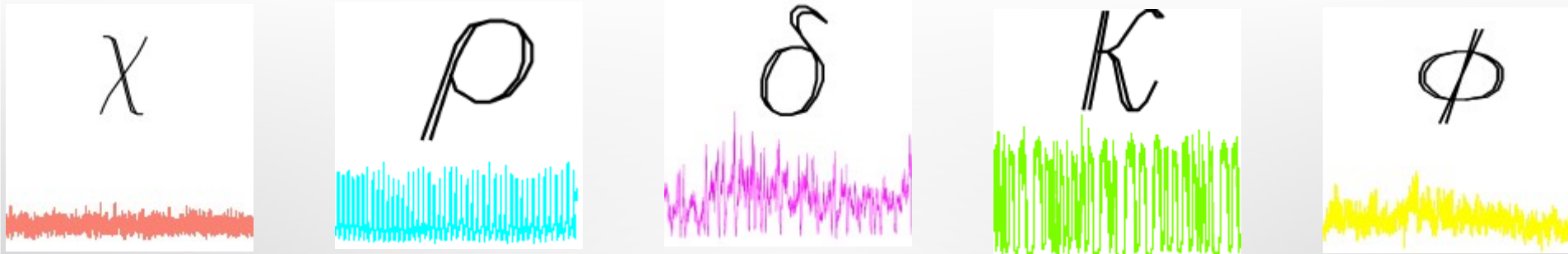
We propose coupling



Results

Application to different states of an X-ray binary (RXTE)

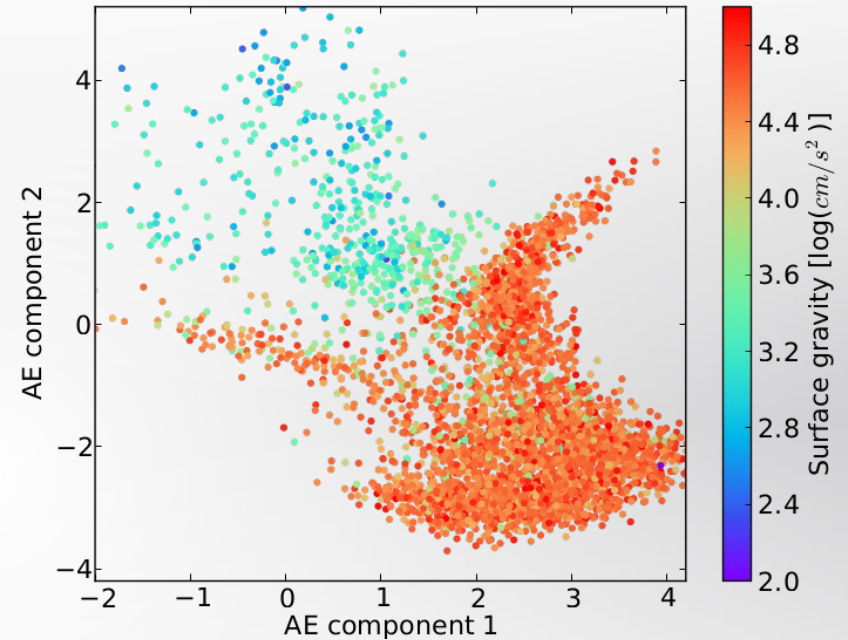
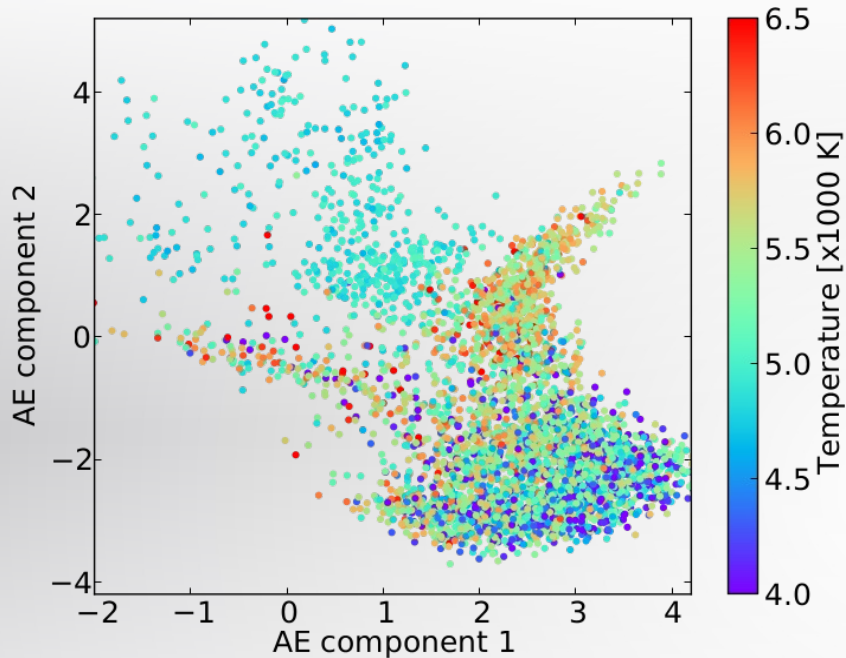
Credit: Harikrishnan et al., RAA, 2012



Gianniotis et al., 2015
submitted to Neurocomputing

Results

Application to Kepler Light Curves



Kügler et al., 2015
submitted to MNRAS

Summary

Irregularly sampled time series

Method

- density-based representation
- capturing all uncertainties
- Inclusion of non-detections
- Avoiding survey specific biases

Results

- Comparable Performance to features
- Omitting arbitrary feature selection
- Allowing for unsupervised methodology
- Exploring transfer learning scheme

Regularly sampled time series

Method

- ESN as vector representation
- Respecting sequential nature
- Coupling visualization to ESN

Results

- Visualization according to latent dynamics
- Recovery of physical properties
- No comparable method exists

Thank you for your attention!



Credit: kepler.nasa.gov