A Neural Network Approach to **Visualising Astronomical Time Series**

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HITS



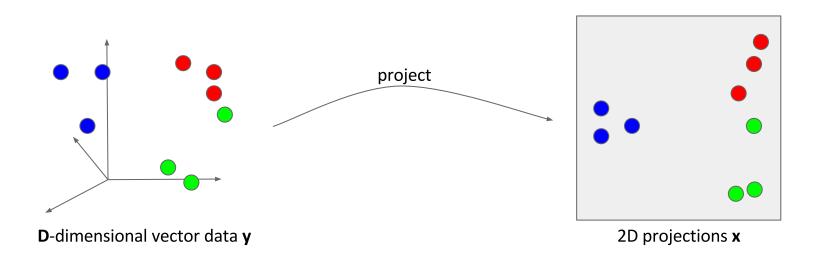


Problem statement

We are working with datasets of time series, and what we would like to do is reduce the dimensionality of the time series.

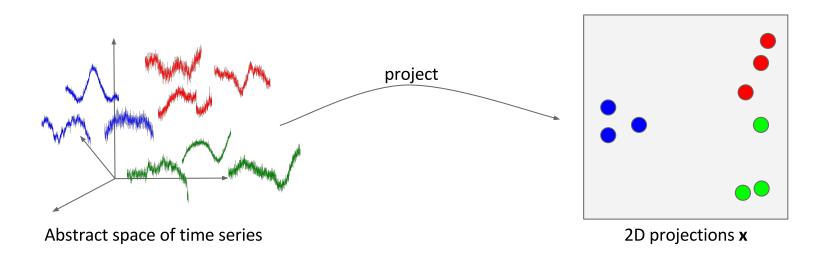
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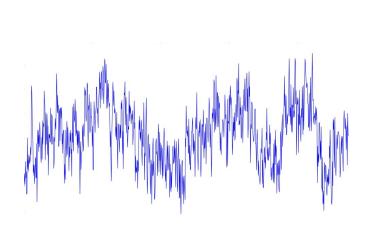
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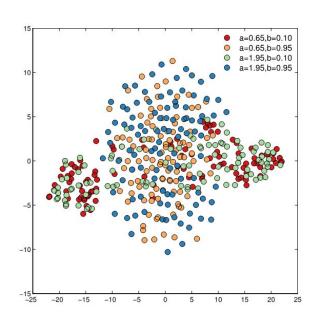
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Is it serious?







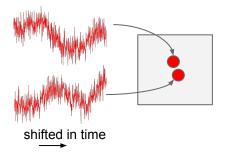
Sequences from Gaussian process with correlation function given by

$$c(x_t, x_{t+1}) = (1 + |h|^{\alpha})^{-\frac{\alpha}{b}}$$





When projecting time series we must account for

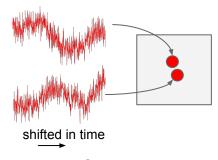


translation invariance

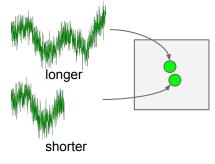




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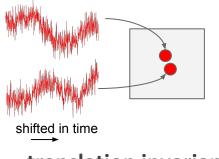


variable length

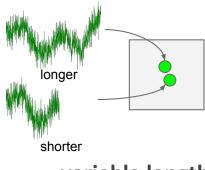
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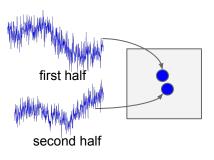
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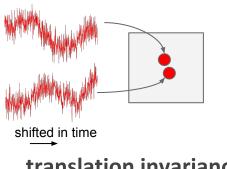


partially observed

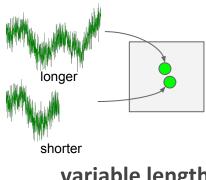




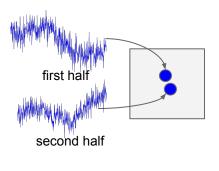
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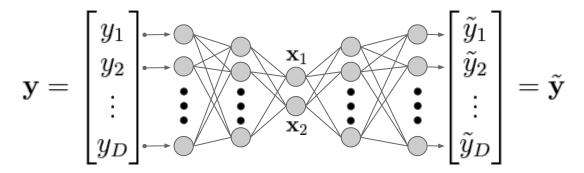
Proposed solution

- come up with a new representation for time series
- reduce dimensionality of new representation via autoencoder



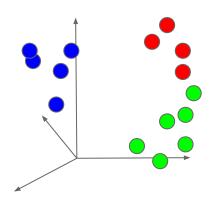


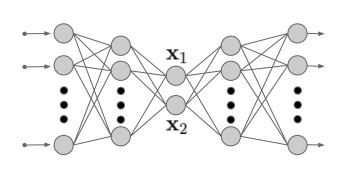
Fan-in fan-out neural network

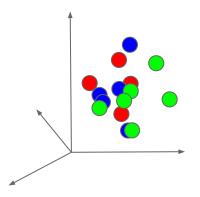


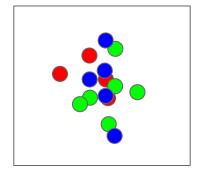


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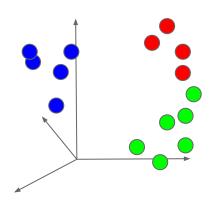


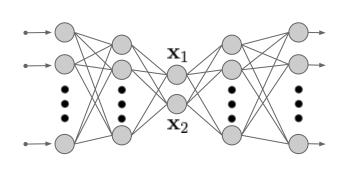


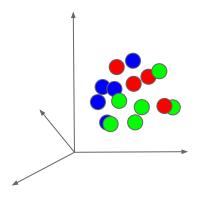
$$\|\mathbf{y} - \tilde{\mathbf{y}}\|^2$$

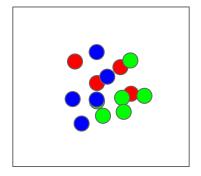


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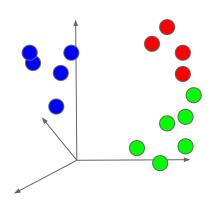


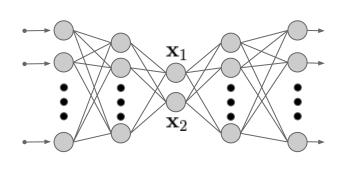
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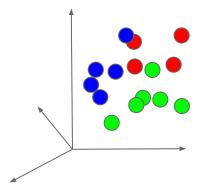


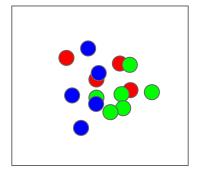


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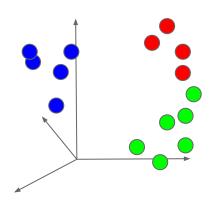


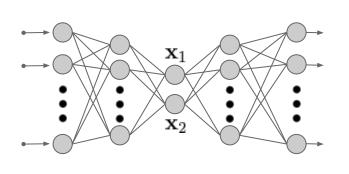


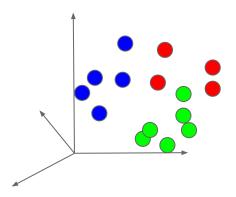
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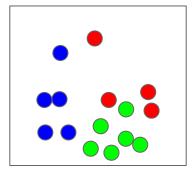


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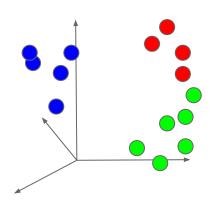


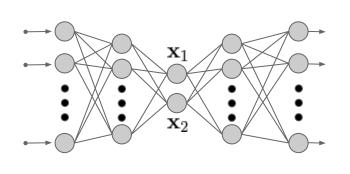


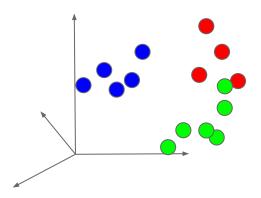
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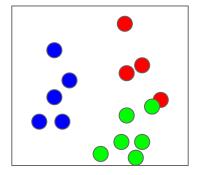


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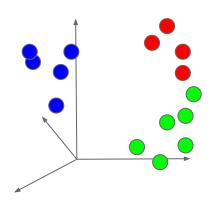


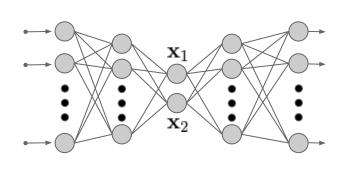


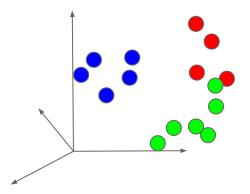
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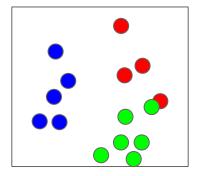


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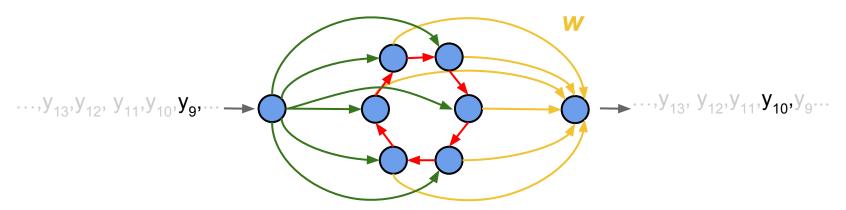


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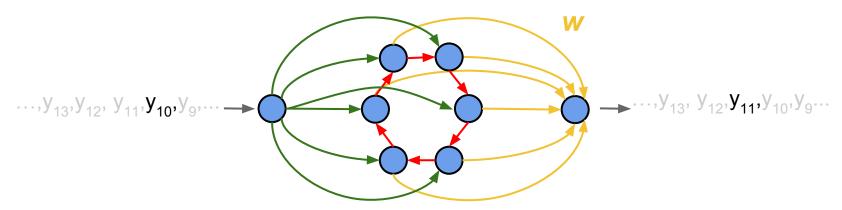
New representation for time series



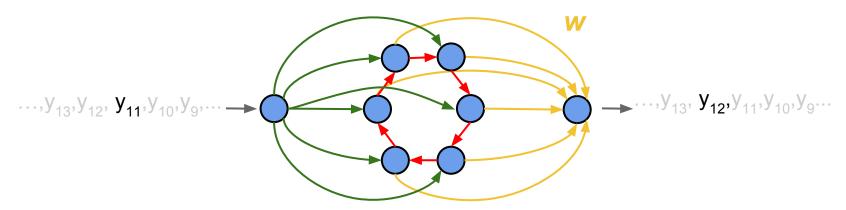
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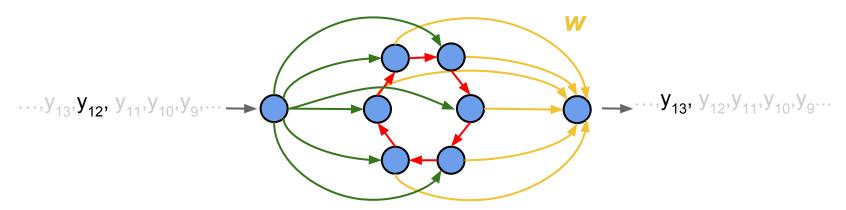
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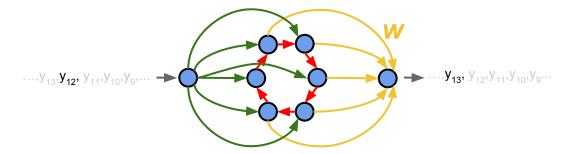


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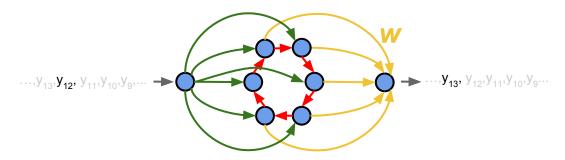
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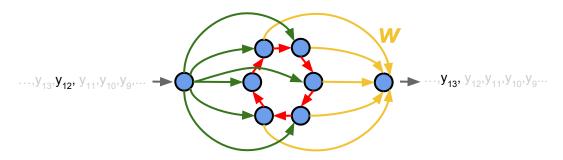


Measure "one-step aheadedness" via g(y;w)

Vector ${f w}$ captures temporal behaviour of ${f y}$

New representation for time series





Measure "one-step aheadedness" via g(y;w)

Vector ${f w}$ captures temporal behaviour of ${f y}$

Fit an ESN model to each time series and obtain

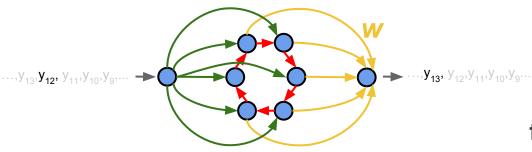
 $\mathbf{y_1} \; o \; \mathbf{w_1}$

 $\mathbf{y_2} \rightarrow \mathbf{w_2}$

 $\mathbf{y_N} \, o \, \mathbf{w_N}$

New representation for time series





Measure "one-step aheadedness" via g(y;w)

Vector \mathbf{W} captures temporal behaviour of \mathbf{y}

Fit an ESN model to each time series and obtain

fitted w invariant to shift, length, partial observation

 $\mathbf{y_1} \; o \; \mathbf{w_1}$

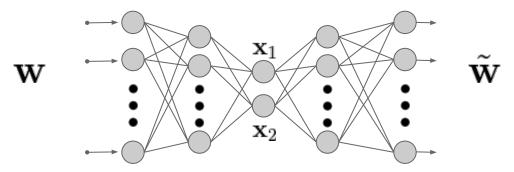
 $\mathbf{y_2} \rightarrow \mathbf{w_2}$

 $ightarrow \mathbf{w}$





Let us autoencode **w** instead of **y**:

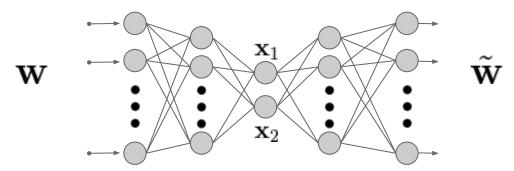


Now we are compressing ESN weight vectors W

Autoencoding of new representation



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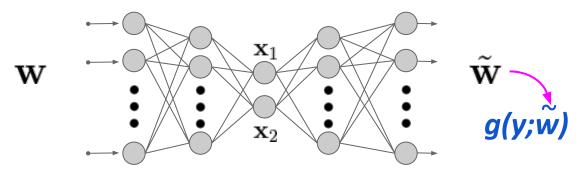


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- lacksquare Don't measure reconstruction with $\|\mathbf{w} ilde{\mathbf{w}}\|^2$



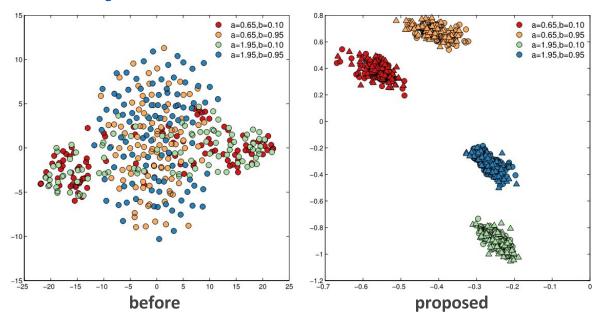


Let us autoencode **w** instead of **y**:



- Now we are compressing ESN weight vectors W
- Don't measure reconstruction with $\|\mathbf{w}\|^2$
- Reconstruction: plug $ilde{\mathbf{W}}$ into g and check how well it still predicts on $extbf{\emph{y}}$

Revisit Cauchy series

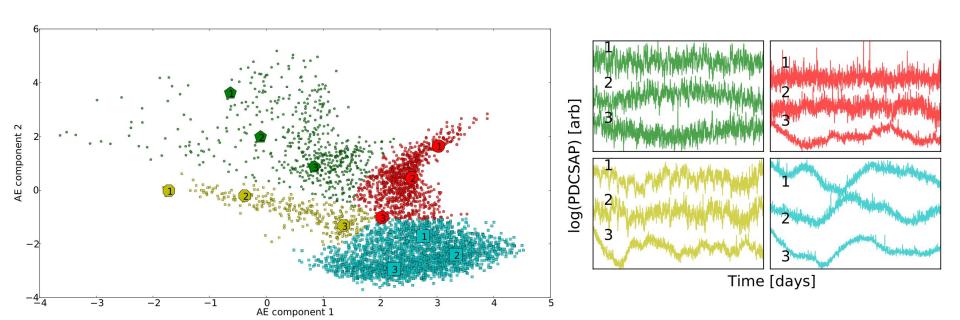


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Real data - Kepler light curves (1)



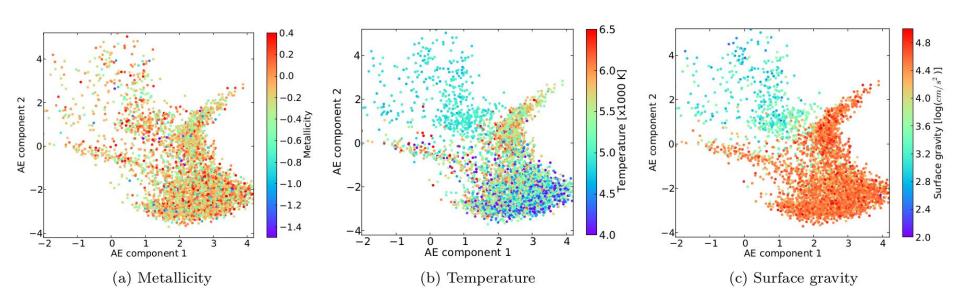


Data taken from online repository of Kepler mission





Physical properties



Surface gravity correlates strongly with variability behaviour

Conclusion



- Time series are qualitatively different entities than vectors
- Latent regime must be accounted in dimensionality reduction
- Currently working on not regularly sampled time series
- For more information please refer to:

Model-Coupled Autoencoder for Time Series Visualisation, Neurocomputing

An Explorative Approach for Inspecting Kepler Data, MNRAS

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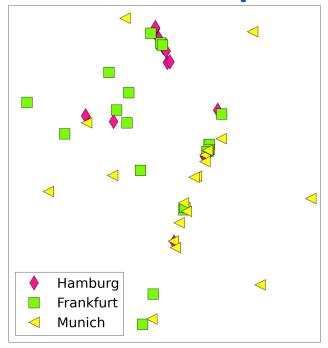
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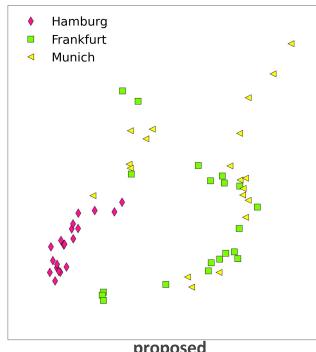
Model-Coupled Autoencoder for Time Series Visualisation, Neurocomputing

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Thank you for your attention!

Real data - Wind speed data



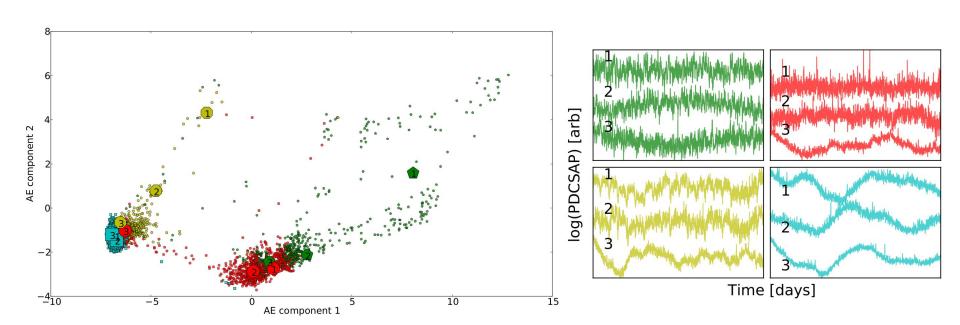


proposed

Data taken from 10 stations around Hamburg, Frankfurt and Munich **Courtesy of Deutscher Wetterdienst**

Real data - Kepler light curves (2)





Time series autoencoded as vectors



Autoencoding of new representation

Let us a

Why is Euclidean distance on parameters meaningless?

- some components w_i have no effect
- some components w_i more sensitive
- some components in different scale
- . . .

Bad idea!

- Now
- lacksquare Don't measure reconstruction with $\|\mathbf{w} ilde{\mathbf{w}}\|^2$





Fan-in fan-out architecture

$$\mathbf{y} = \begin{bmatrix} y_1 \\ y_2 \\ \vdots \\ y_D \end{bmatrix} \xrightarrow{\mathbf{x}_1} \begin{bmatrix} \tilde{y}_1 \\ \tilde{y}_2 \\ \vdots \\ \tilde{y}_D \end{bmatrix} = \tilde{\mathbf{y}}$$

$$f_{enc}(\mathbf{y}): \mathbb{R}^D \to \mathbb{R}^2 \quad f_{dec}(\mathbf{x}): \mathbb{R}^2 \to \mathbb{R}^D$$