

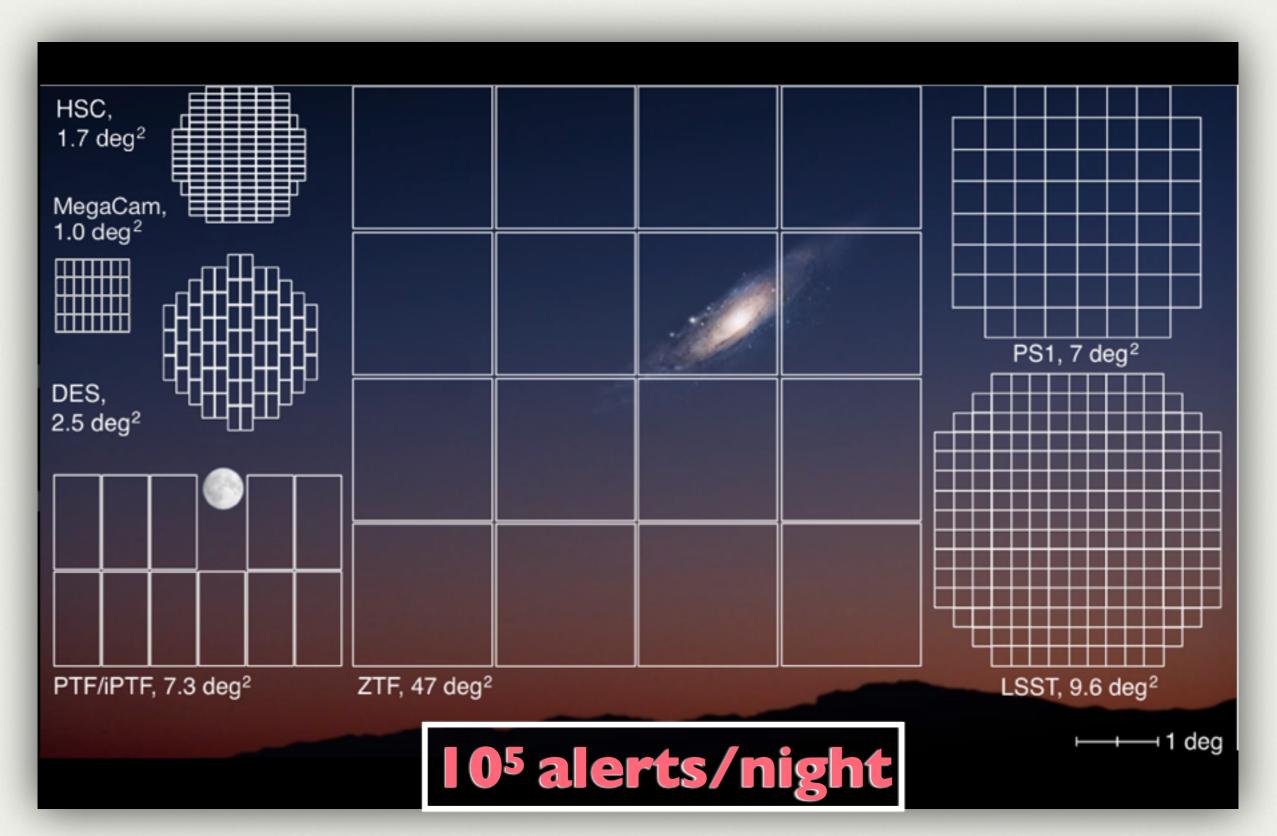
AMPEL

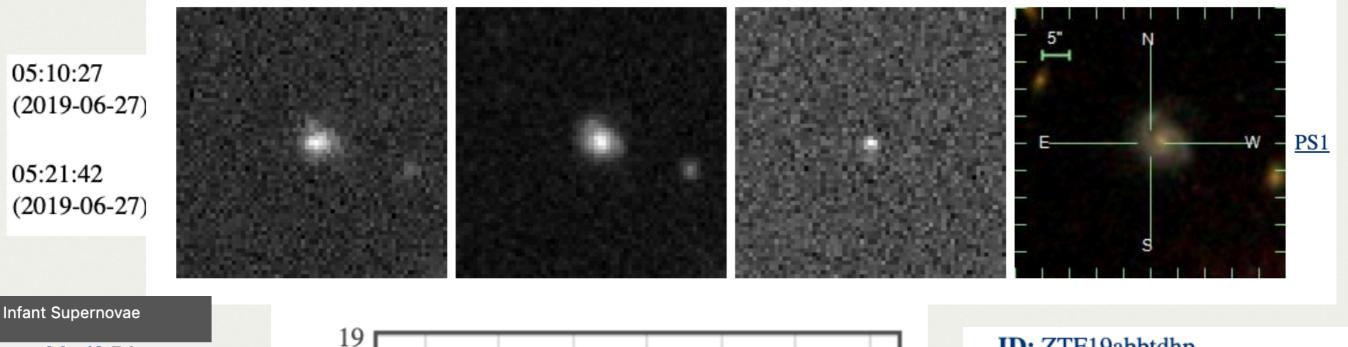
Valéry Brinnel Humboldt Universität zu Berlin

AG 2019

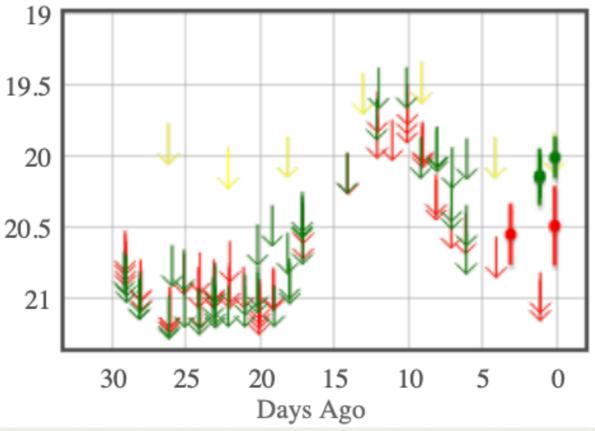
INTRODUCTION

MOTIVATION

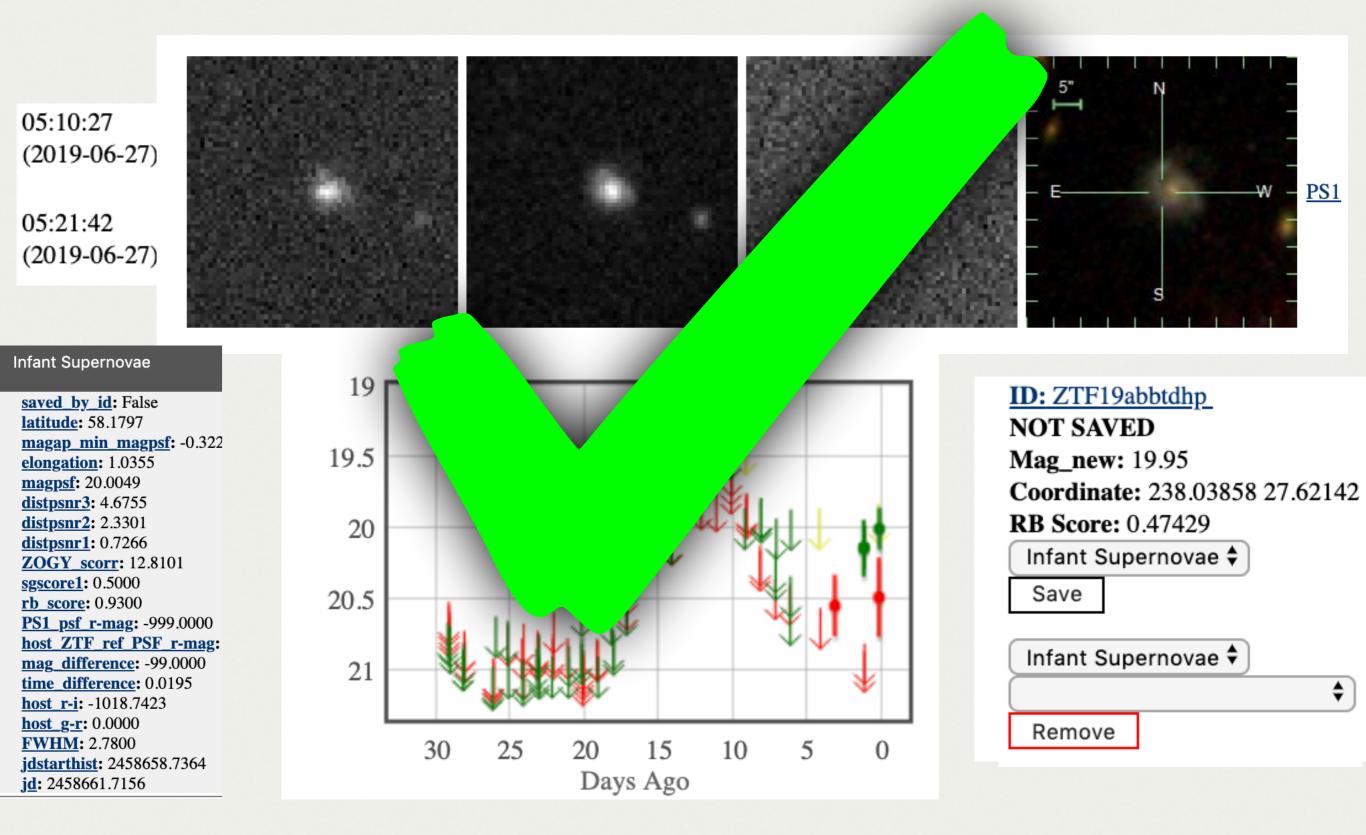


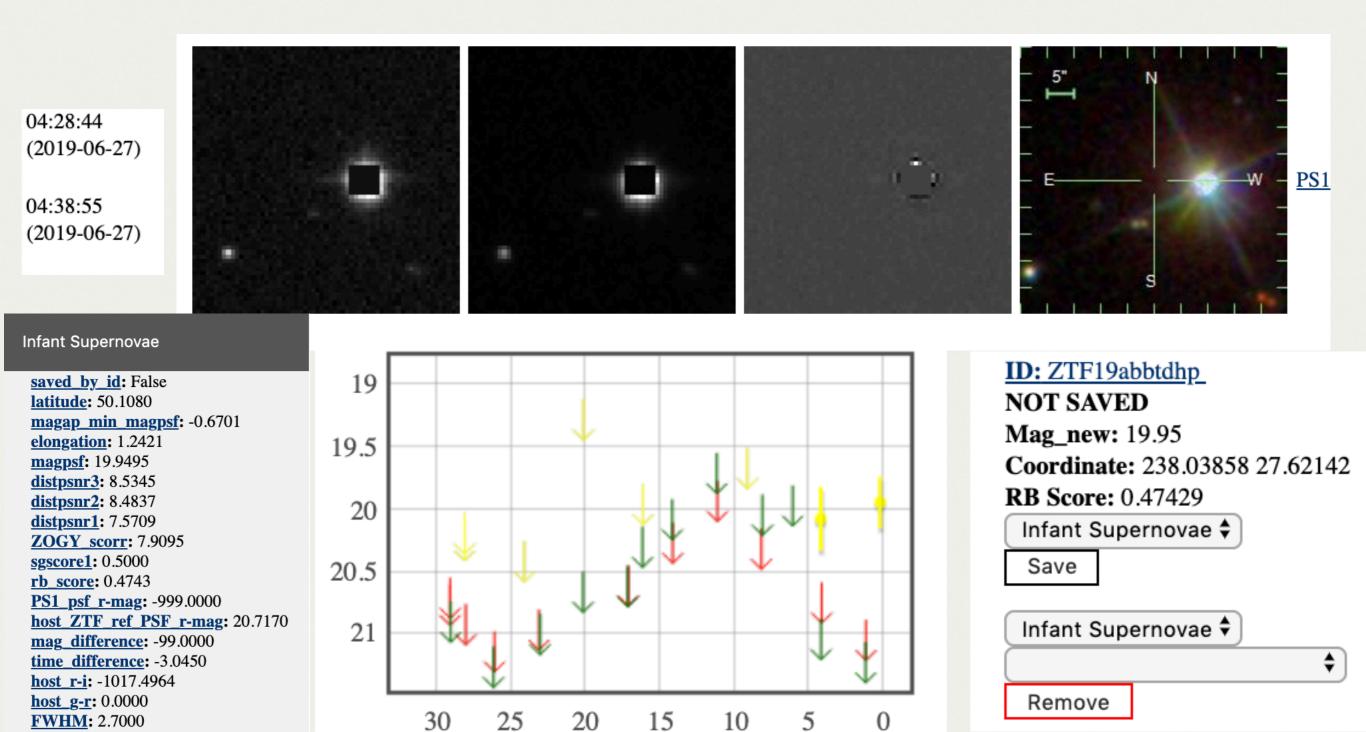


saved_by_id: False latitude: 58.1797 magap_min_magpsf: -0.322 elongation: 1.0355 magpsf: 20.0049 distpsnr3: 4.6755 distpsnr2: 2.3301 distpsnr1: 0.7266 **ZOGY_scorr:** 12.8101 sgscore1: 0.5000 **rb_score:** 0.9300 **PS1_psf_r-mag:** -999.0000 host_ZTF_ref_PSF_r-mag: mag difference: -99.0000 time difference: 0.0195 host_r-i: -1018.7423 host_g-r: 0.0000 **FWHM:** 2.7800 jdstarthist: 2458658.7364 jd: 2458661.7156









Days Ago

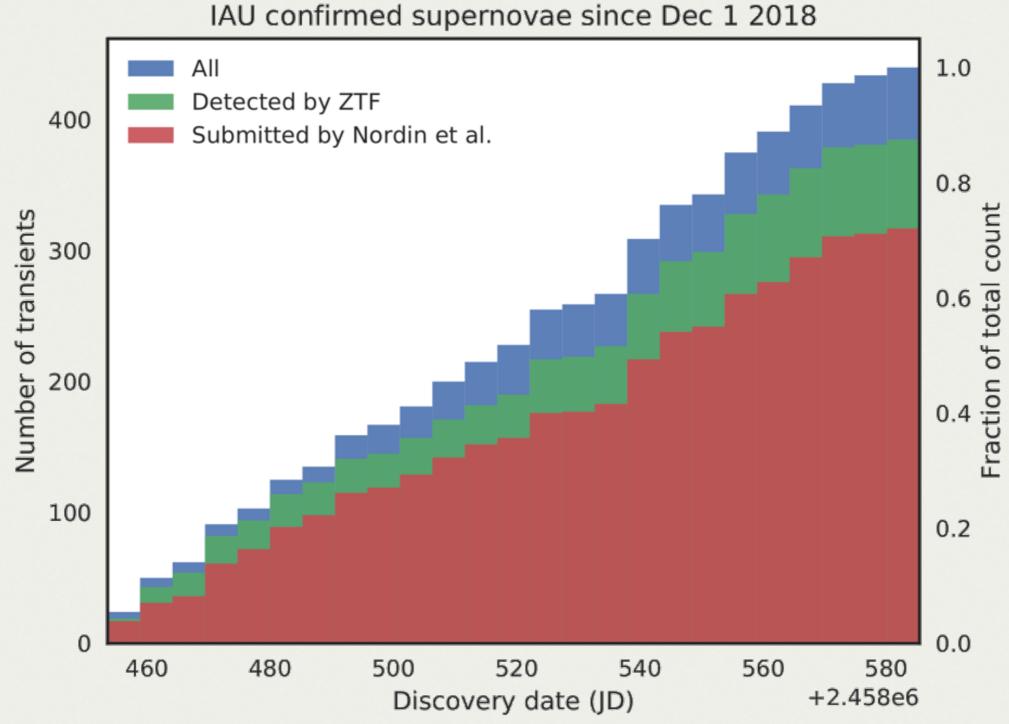
jdstarthist: 2458657.6869 jd: 2458661.6866



MOTIVATION

• Automatisation

AMPEL NOW



Nordin et al. 2019

MOTIVATION

- Automatisation
- Multi-messenger
 - Neutrino
 - GW

AMPEL NOW

TITLE: GCN CIRCULAR NUMBER: 25616 SUBJECT: LIGO/Virgo S190901ap: Candidates from the Zwicky Transient Facility DATE: 19/09/02 14:32:28 GMT FROM: Mansi M. Kasliwal at Caltech/Carnegie <mansikasliwal@gmail.com>

Erik Kool (OKC), Robert Stein (DESY), Yashvi Sharma (Caltech), Viraj Karambelkar (Caltech), Mansi Kasliwal (Caltech), Daniel Perley (LJMU), Valery Brinnel (HU Berlin), Jakob Nordin (HU Berlin), Shreya Anand (Caltech), Michael Coughlin (Caltech), Leo P. Singer (NASA GSFC), Igor Andreoni (Caltech), Gaurav Waratkar (IITB), Harsh Kumar (IITB), Maitreya Khandagale (IITB), Kunal Deshmukh (IITB), Varun Bhalerao (IITB), G. C. Anupama (IIA), Dougal Dobie (USyd/CSIRO), Brad Cenko (NASA GSFC), Tomas Ahmuda (UMD), Eric Bellm (UW), Albert Kong (NTHU), Anna Franckowiak (DESY), Pradip Gatkine (UMD)

On behalf of the Zwicky Transient Facility (ZTF) and Global Relay of Observatories Watching Transients Happen (GROWTH) collaborations

We observed the localization region of the gravitational wave trigger S190901ap (LVC et al. GCN 25606, GCN 25614) with the Palomar 48-inch telescope equipped with the 47 square degree ZTF camera (Bellm et al. 2019, Graham et al. 2019). The tiling was optimally determined and triggered

https://gcn.gsfc.nasa.gov/gcn3/25616.gcn3

AMPEL



Bad transient

Take a closer look

Do something... now!

PEOPLE





L. Rauch



M. Rigault



R. Stein



N. Miranda









S. van Velzen

C. Ward U. Feindt

STRUCTURE

STRUCTURE

- Execution layers (tiers)
- Channels

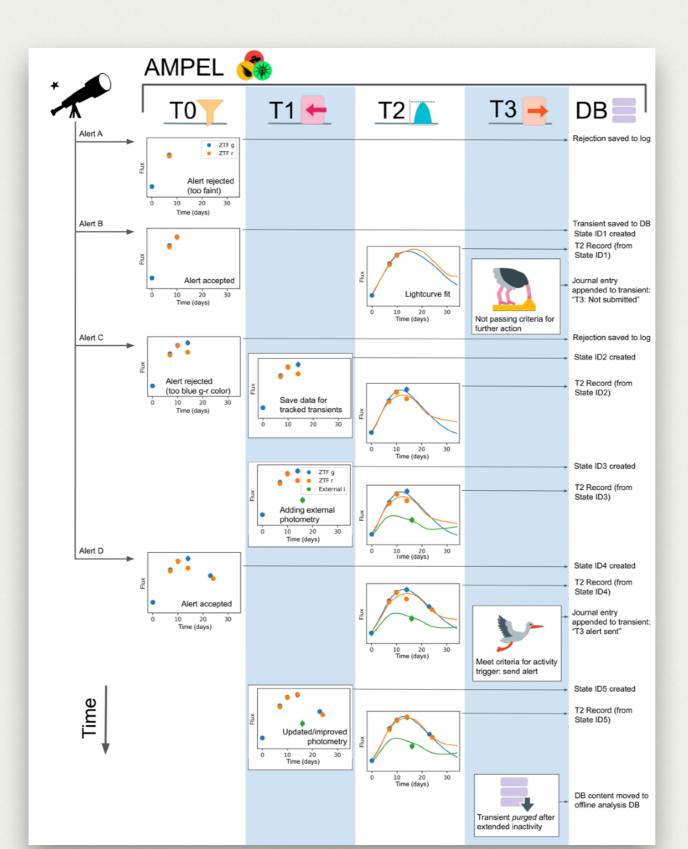
STRUCTURE

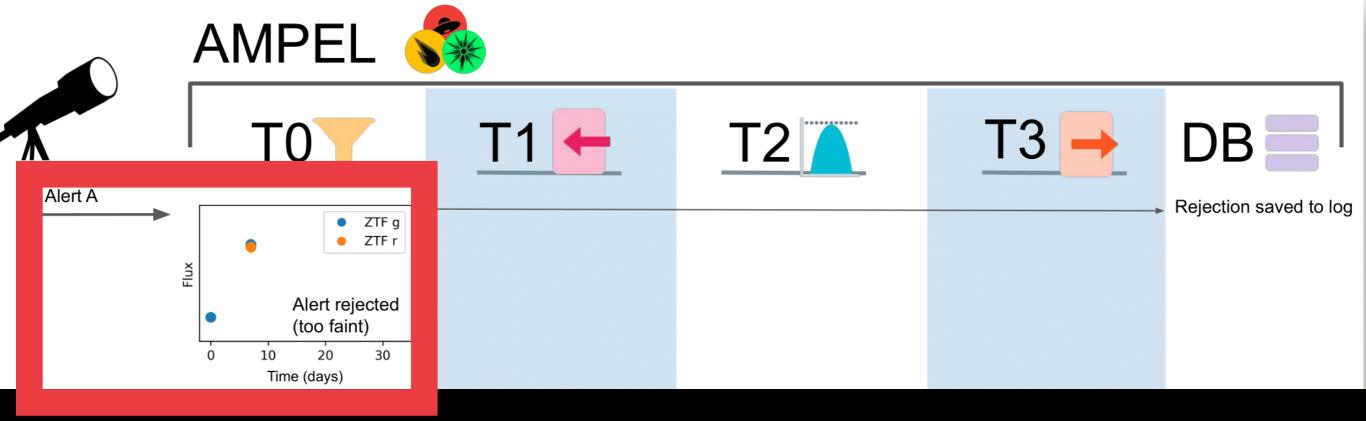
DB & Execution layers

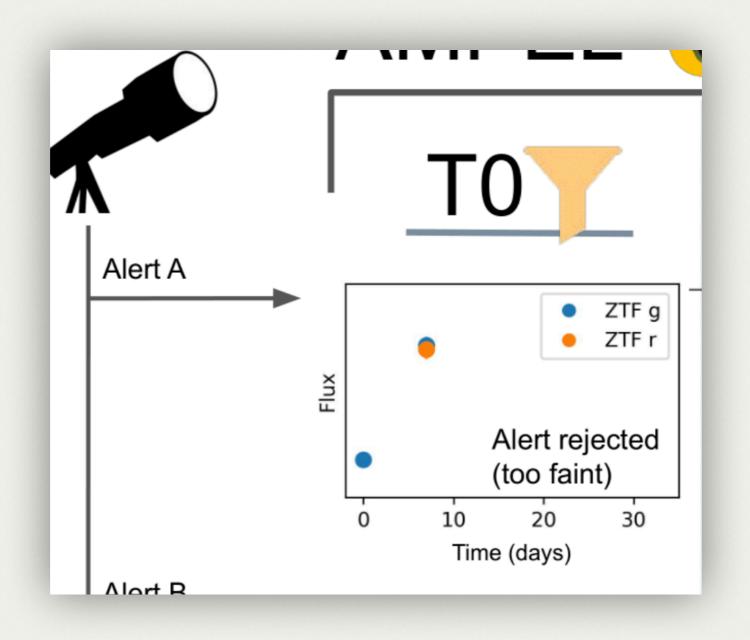
- Ampel has 4 tiers
- Independently scheduled (strictly speaking not a *pipeline*)
- Each tier has a different purpose

| Tier | General | |
|------|---------|--|
| 0 | Add | |
| | Combine | |
| 2 | Augment | |
| 3 | React | |

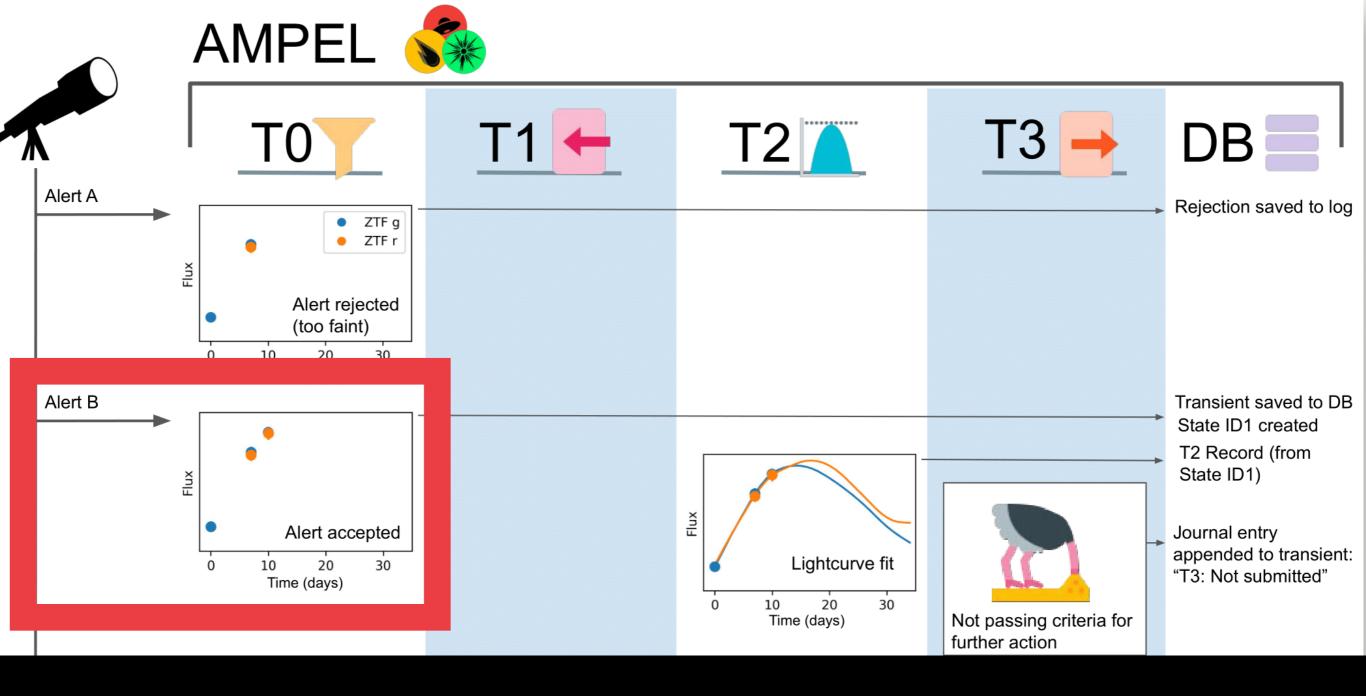
| Tier | General | ZTF |
|------|---------|----------------------|
| 0 | Add | Filter & save alerts |
| | Combine | Create states |
| 2 | Augment | Evaluate Priority |
| 3 | React | React |

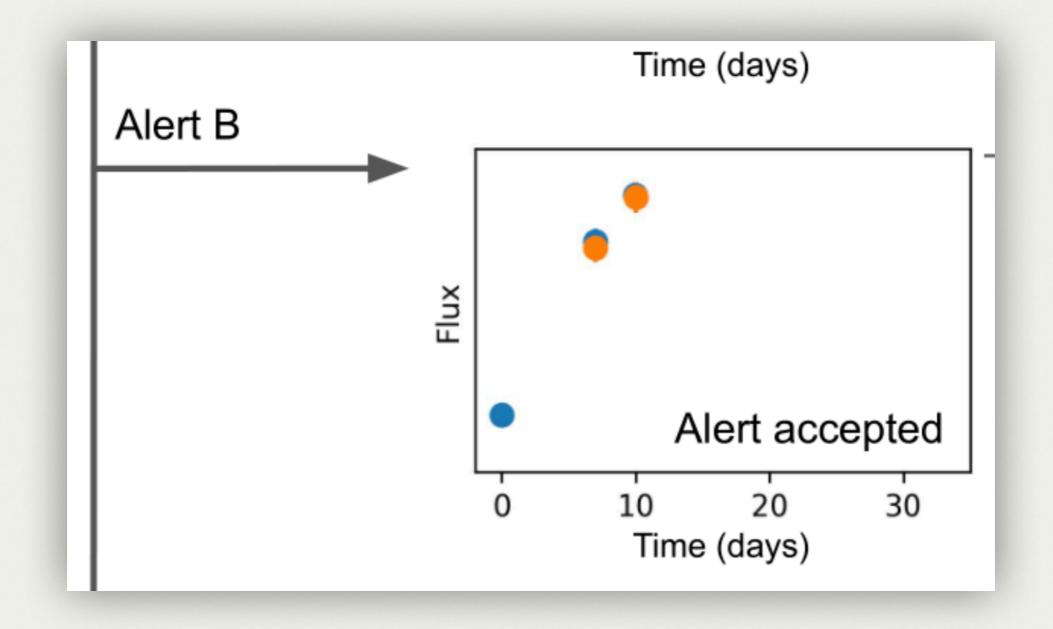




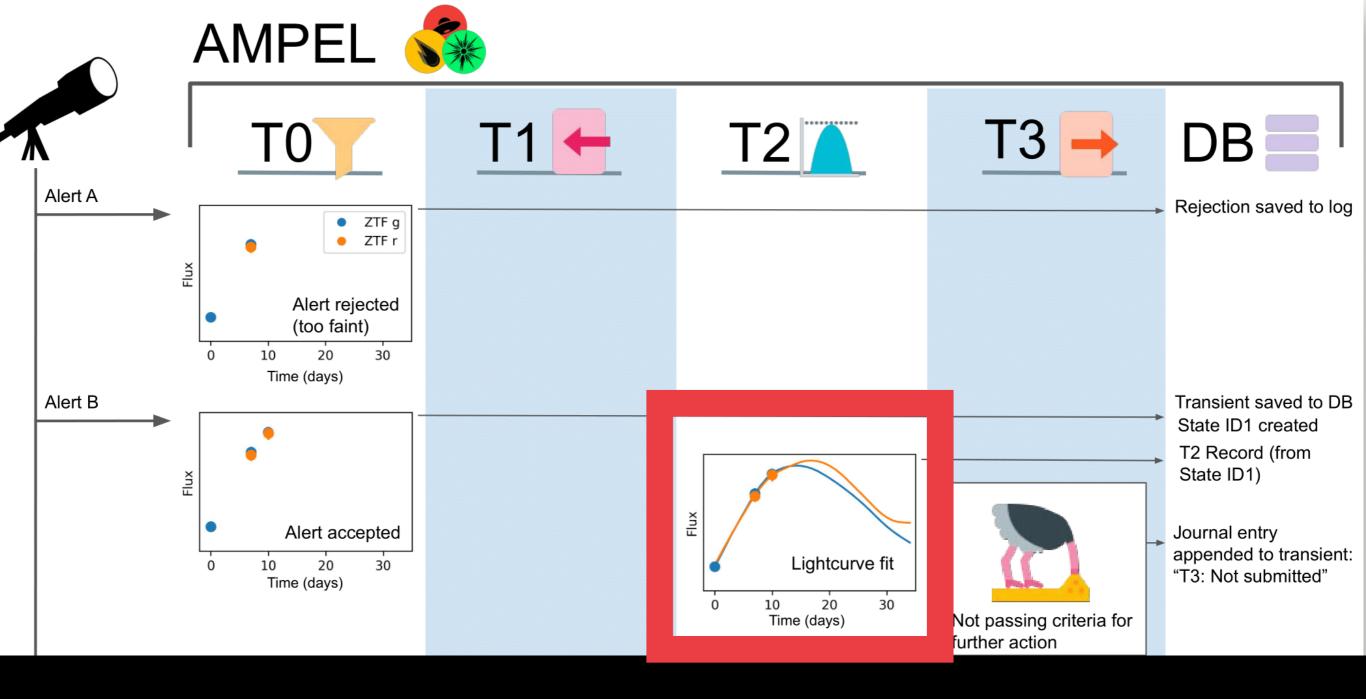


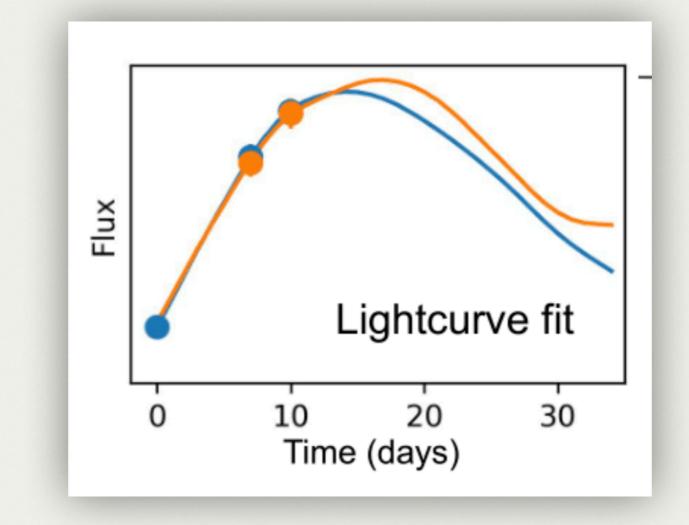
Rejection saved to logs



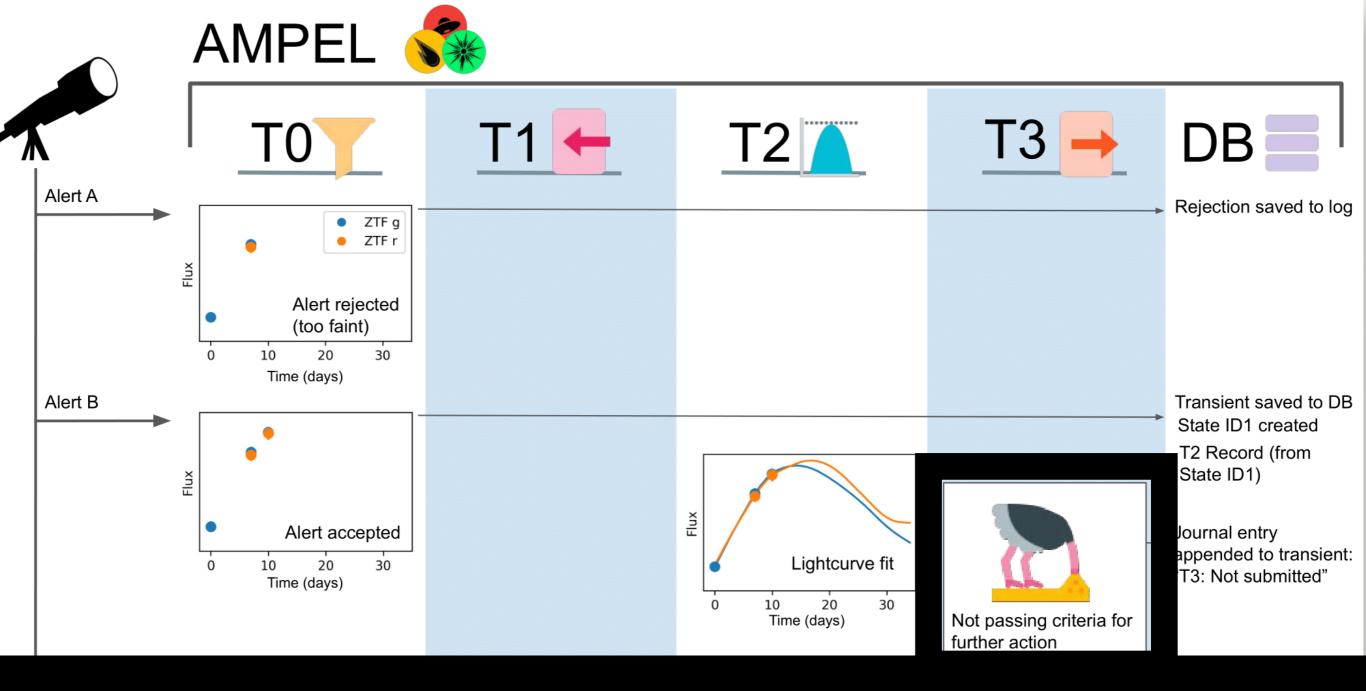


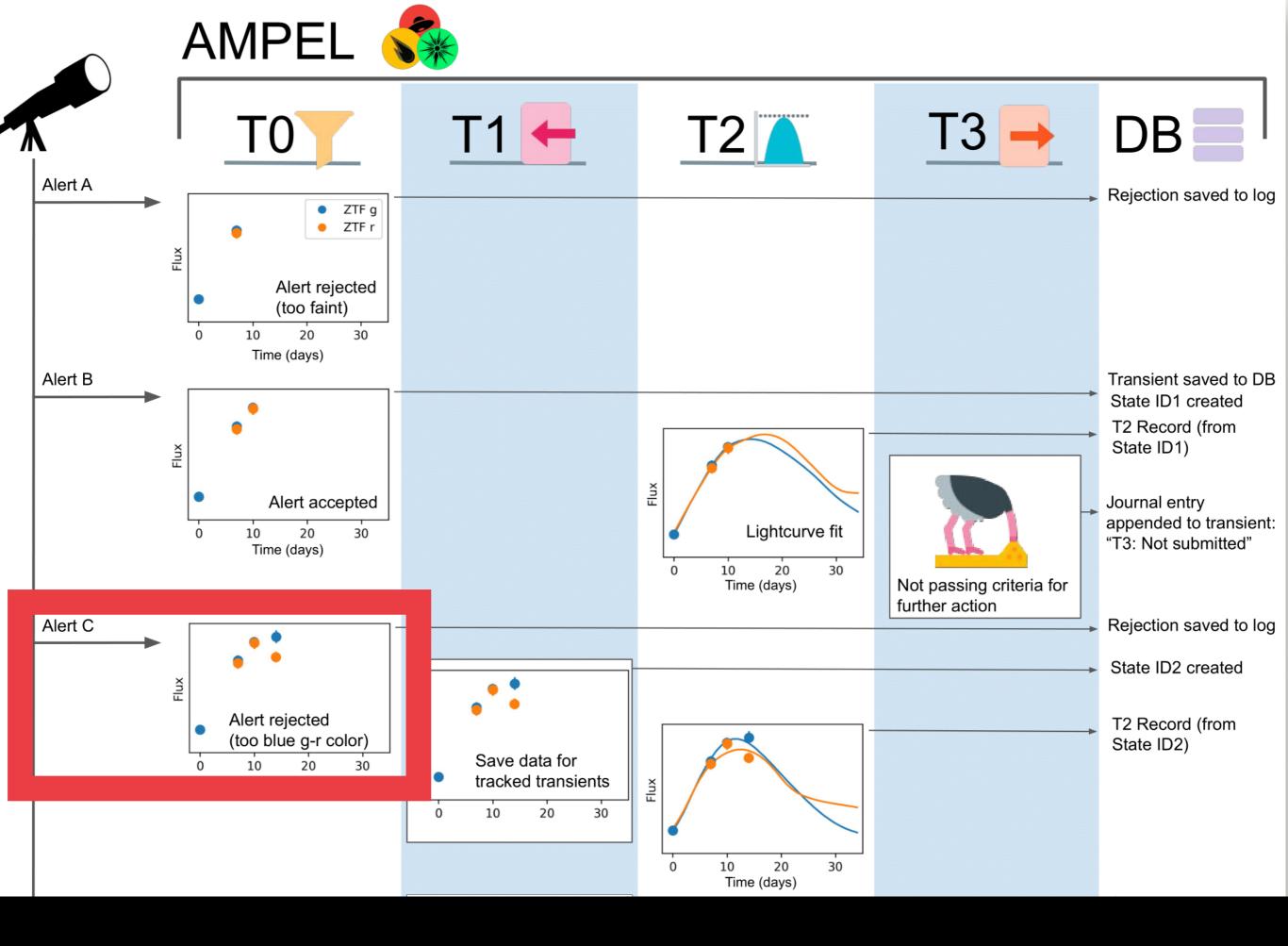
Alert passes filter criteria

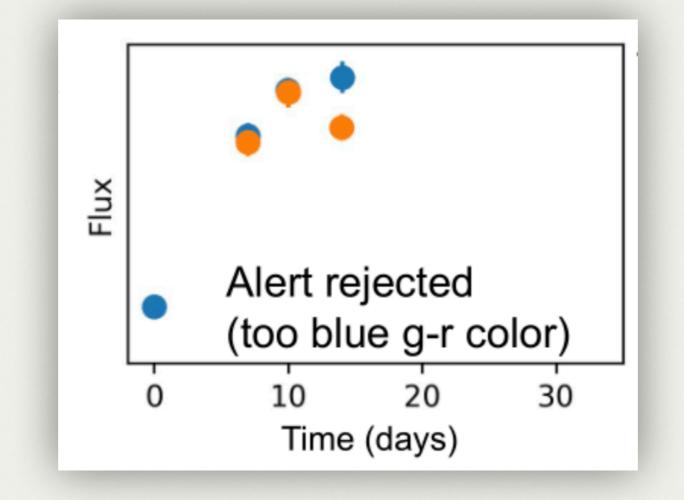




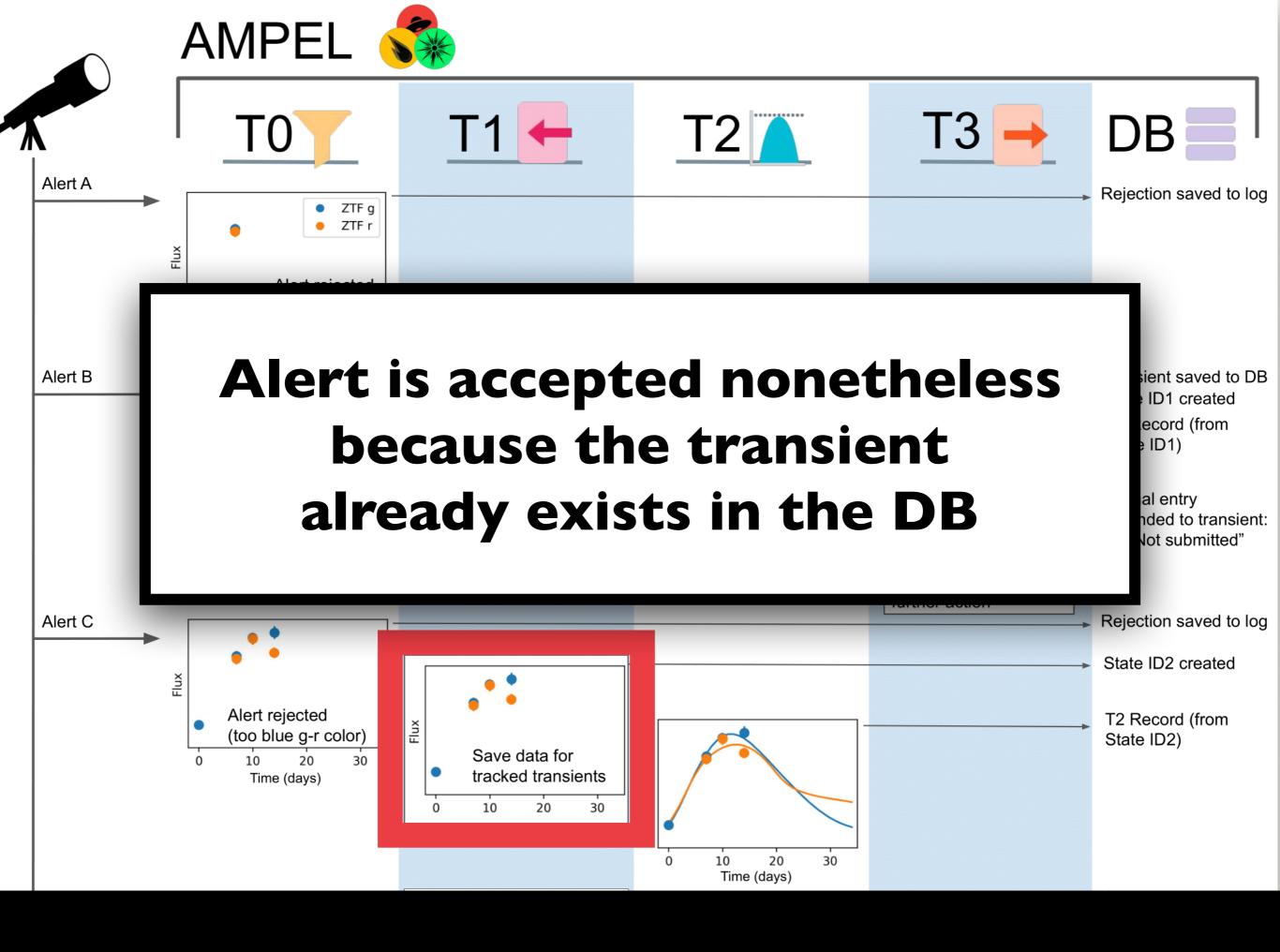
T2 unit SNCOSMO is run

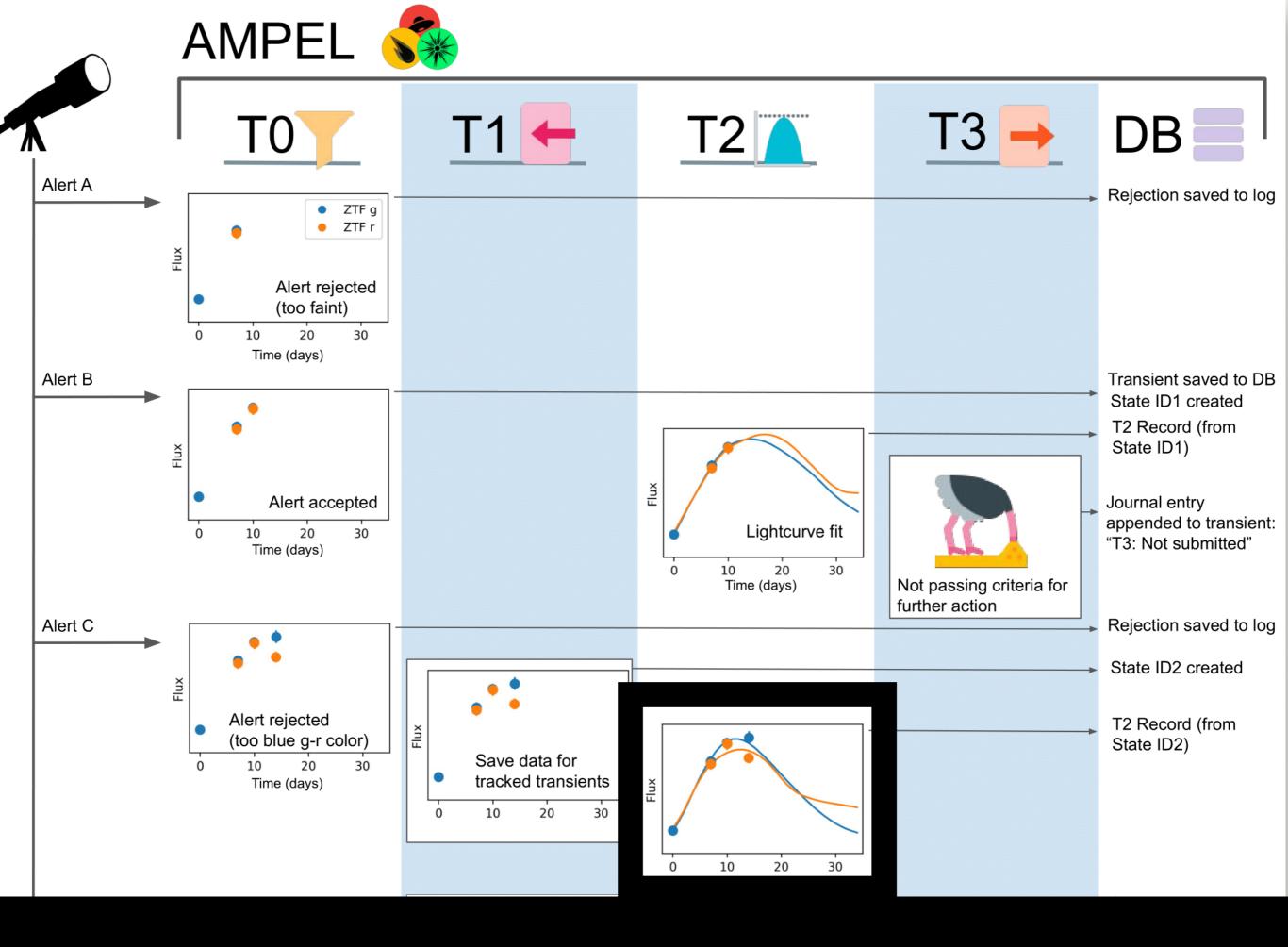




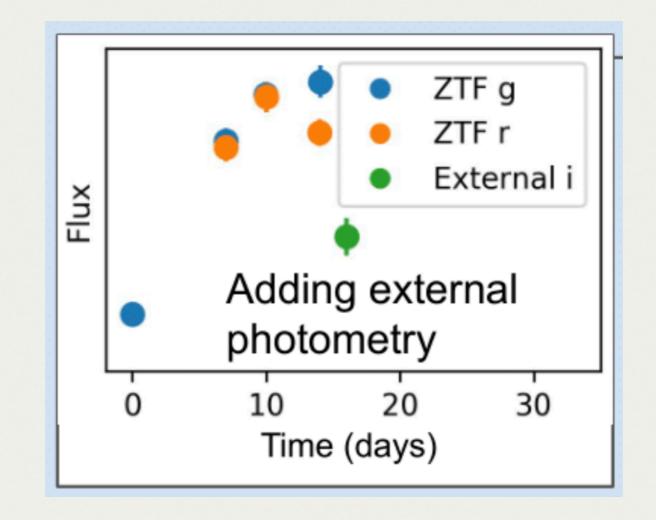


Filter rejects alert but ...



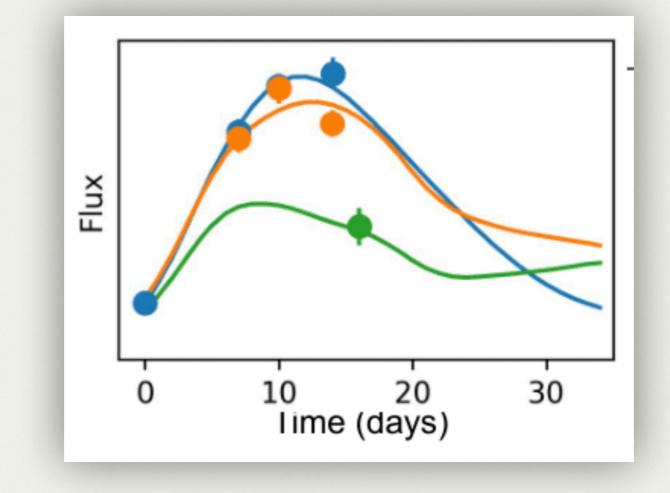






Create new state with external photometry





For which we fit lightcurve templates (T2)



CHANNEL

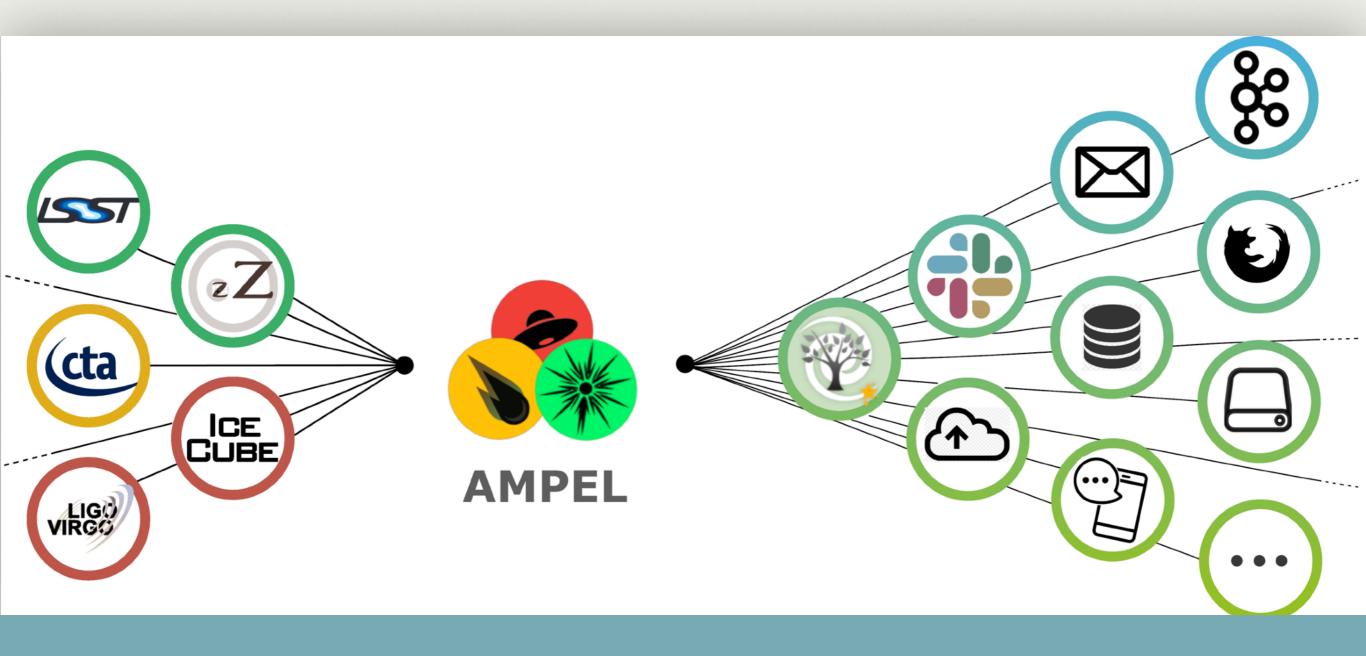
- Sets the following parameters:
 - Which instruments ?
 - Which filter ?
 - Which T2 units ?
 - Which T3 units ?

DEFINE YOUR AMPEL

- How transient(s) should be analysed
- What are your outputs

Ampel applies "analysis schema" consistently to a real-time data

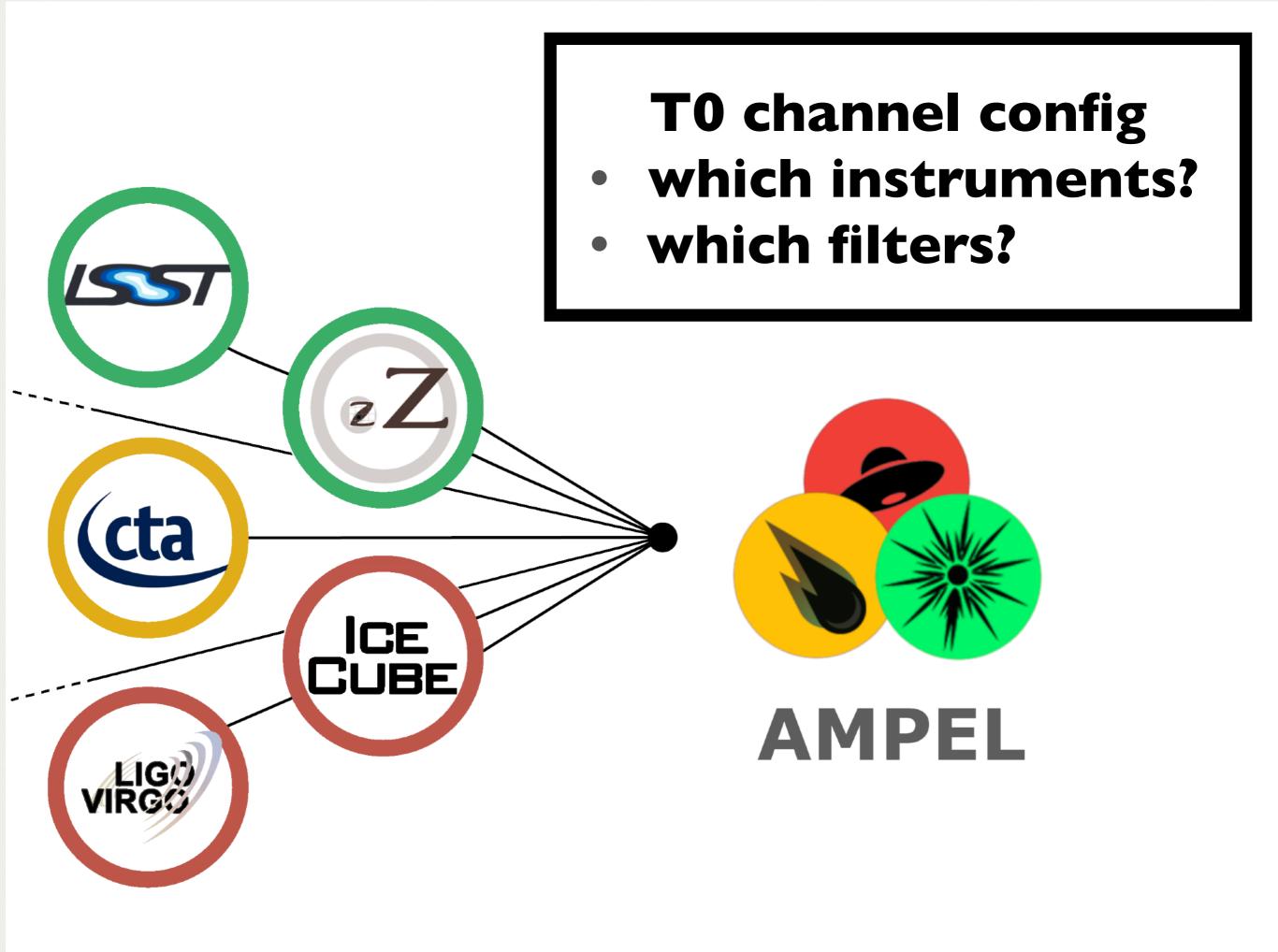
AMPEL

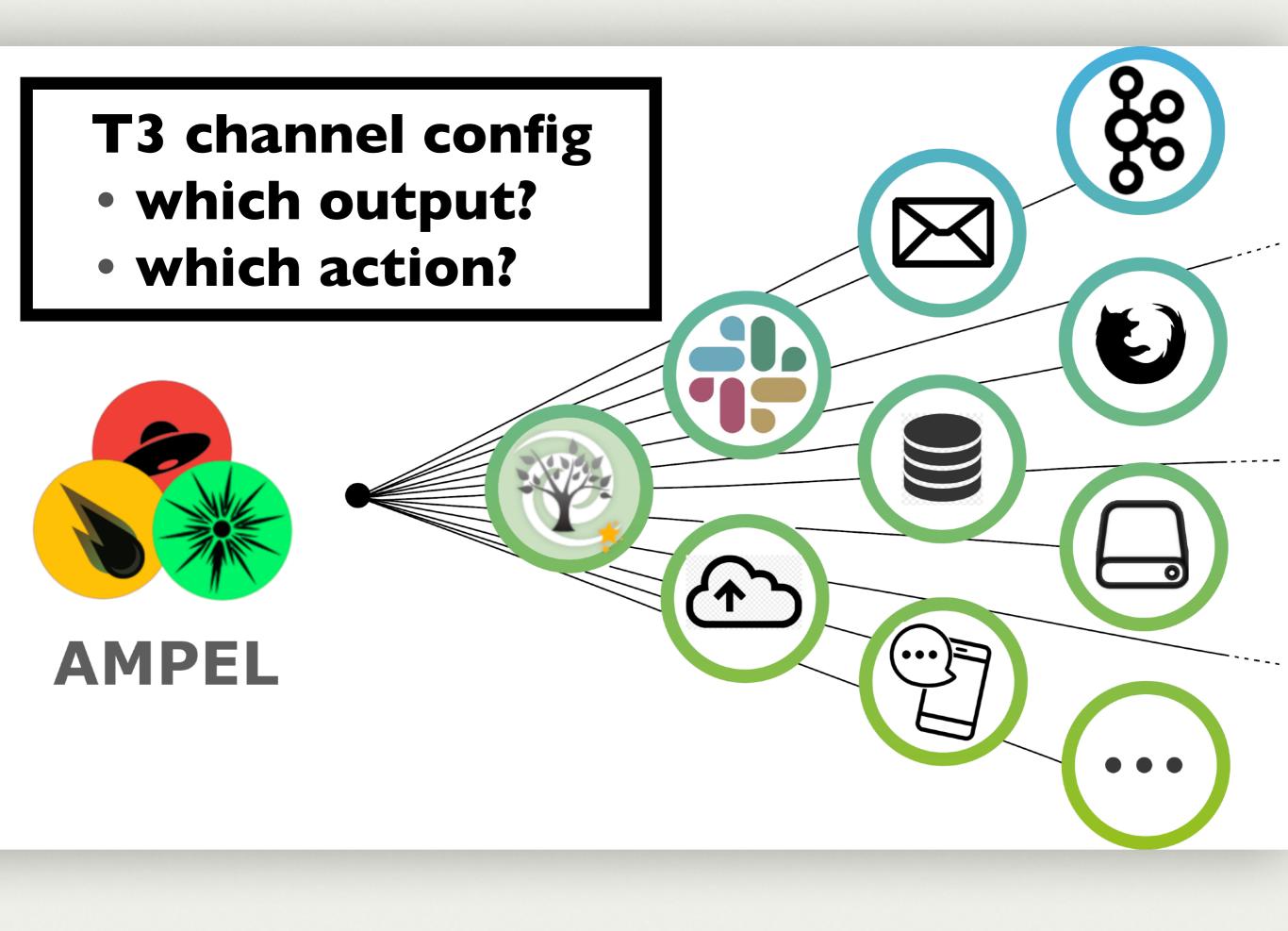


T3

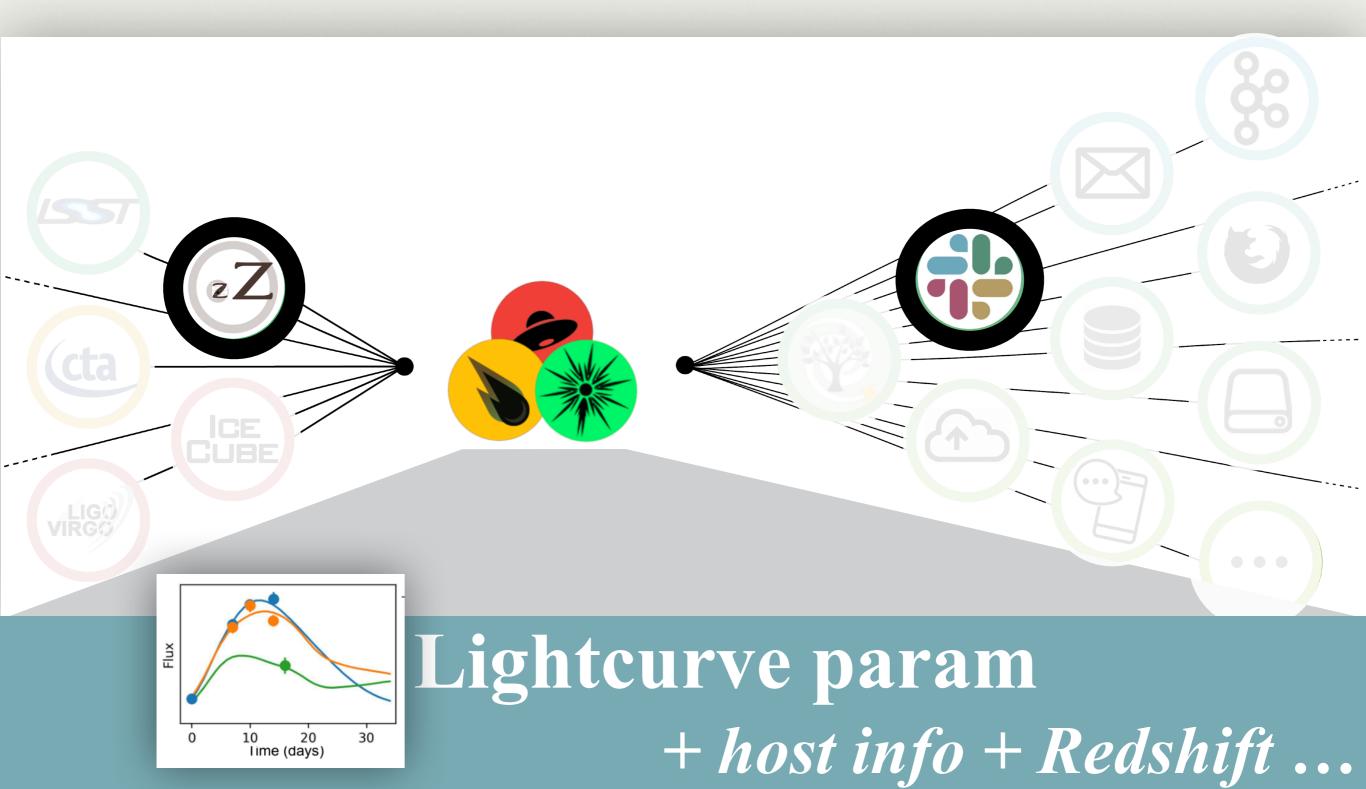
T1/T2

T0





YOUR AMPEL



STRUCTURE

DB & Execution layers

Python implementation

PYTHON IMPLEMENTATION

- Abstract classes regulate input / output
- Ampel units stored in GitHub repositories
- Ampel-core + Ampel-base is ~20K lines of code

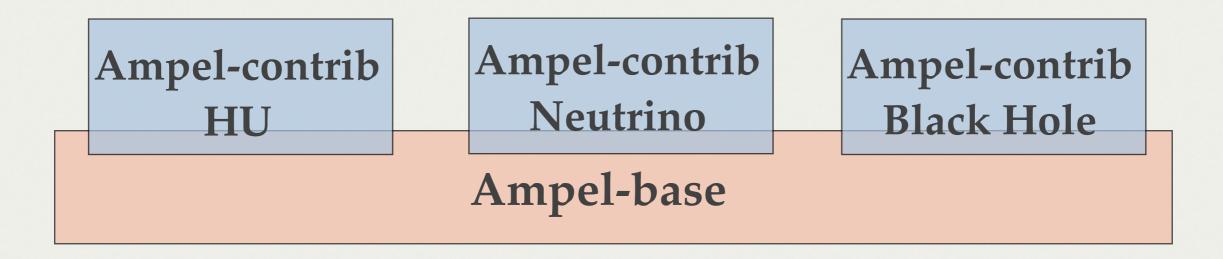
CONTRIBUTED UNITS

| AmpelProject / Ampel-contrib-HU Private | O Unwatch 🗸 | 13 | Star 0 | % Fork 0 |
|--|----------------|----------|------------|------------|
| <> Code ① Issues ④ ⑦ Pull requests ① Projects 0 	 Wiki ① Security II Insig | ghts 🛛 🏠 Setti | ings | | |
| Branch: master - Ampel-contrib-HU / ampel / contrib / hu / t2 / T2SNCosmo.py | | | Find file | Copy path |
| jvansanten T2SNCosmo: back off on errno 24 | | | 41fbf42 | 9 days ago |
| 5 contributors 🔯 📓 🗐 📬 | | | | |
| Execu File 386 lines (321 sloc) 14.3 KB | | Raw Blam | ne History | e 🗊 |
| <pre>1 # 'usr/bin/env python 2 # .*- coding: utf-8 -*-</pre> | | | | |
| | - | - | | - |
| | | | | |

Contributions

Light curve fitter

CODE REPOSITORIES



Internal

Ampel-core

STRUCTURE

DB & Execution layers

Other for a contraction of the contract of the

STRUCTURE

DB & Execution layers

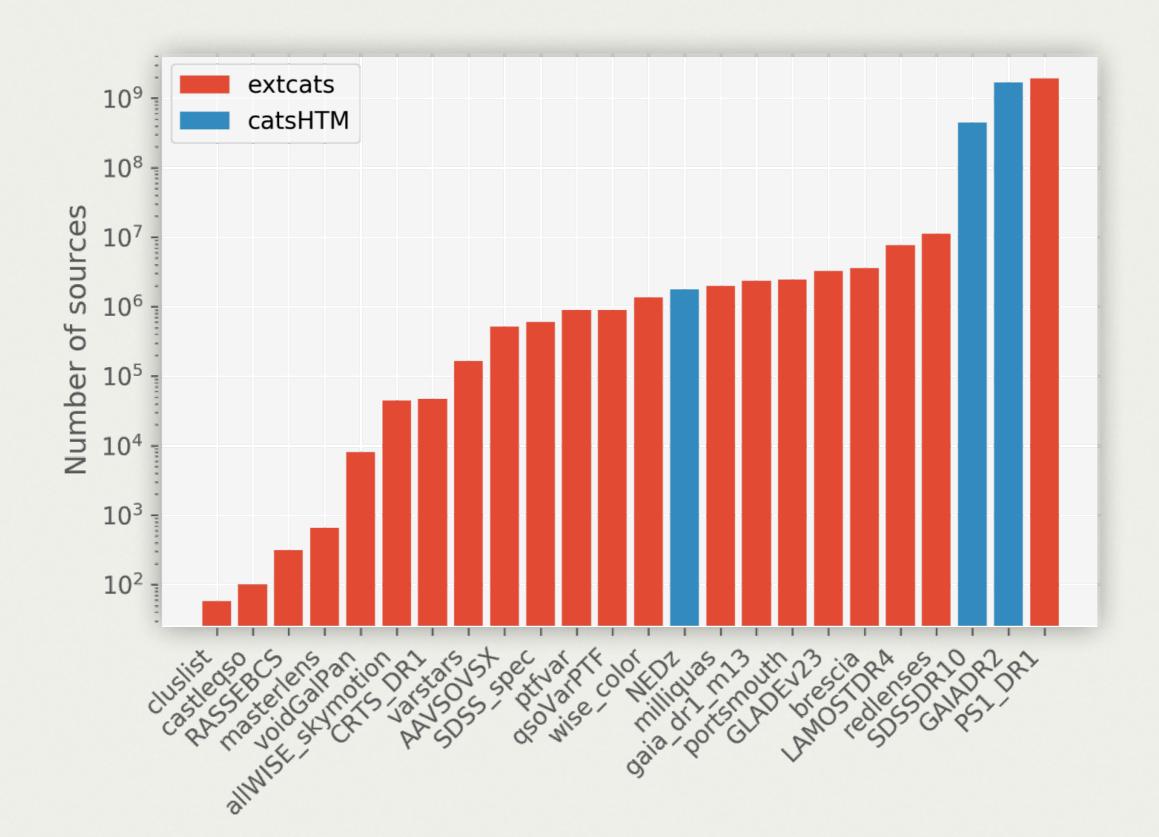
Implementation

Containerization

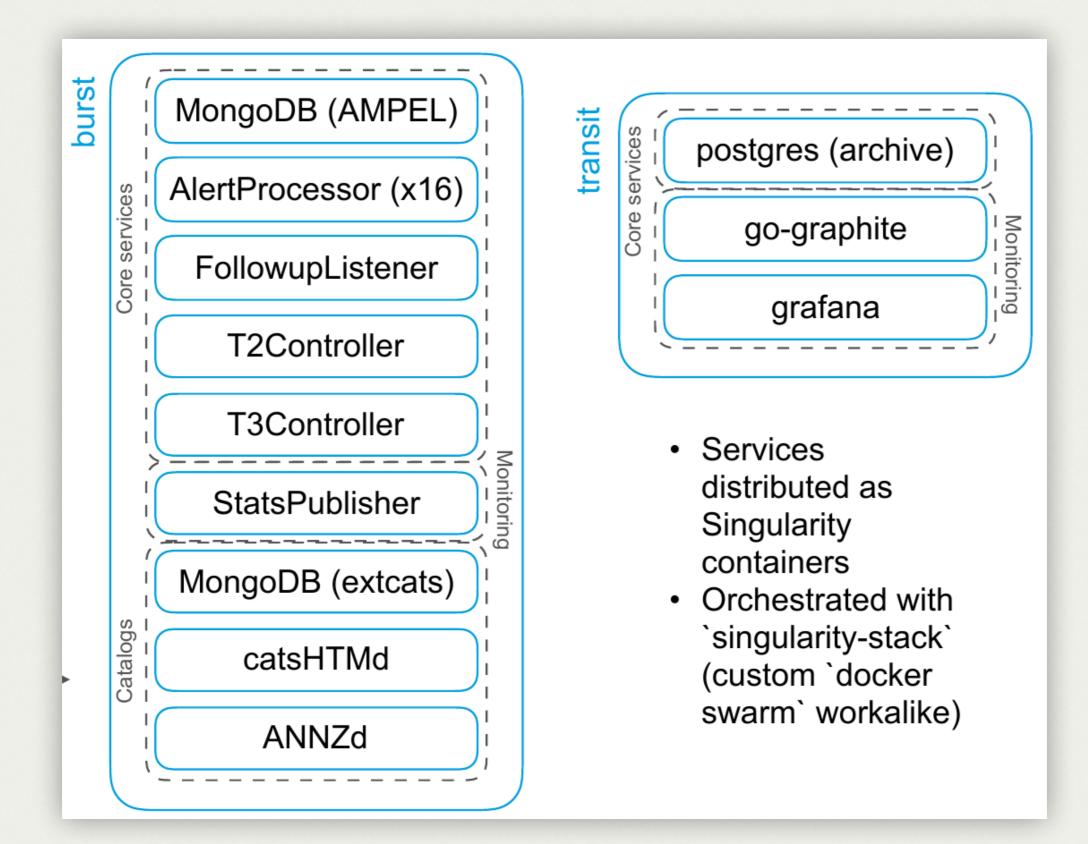


LIVE AMPEL

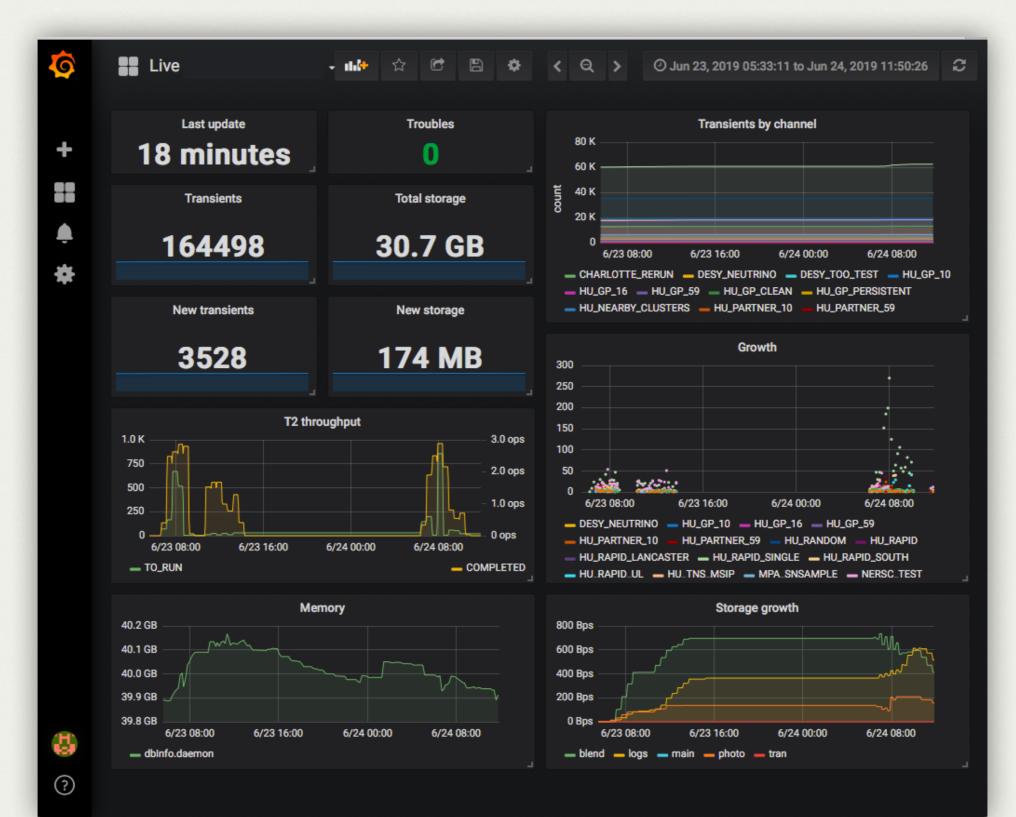
CATALOGS



THE AMPEL STACK



THE AMPEL STACK



WHY BOTHER?





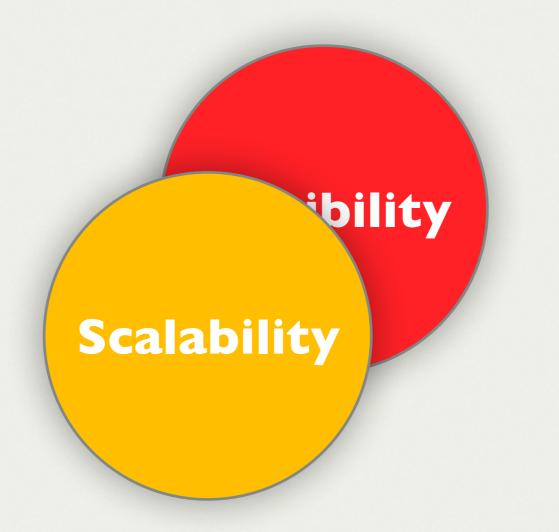
COSTS: TIME

- Understand how AMPEL works
- Break down your analysis
 in adequate independent units

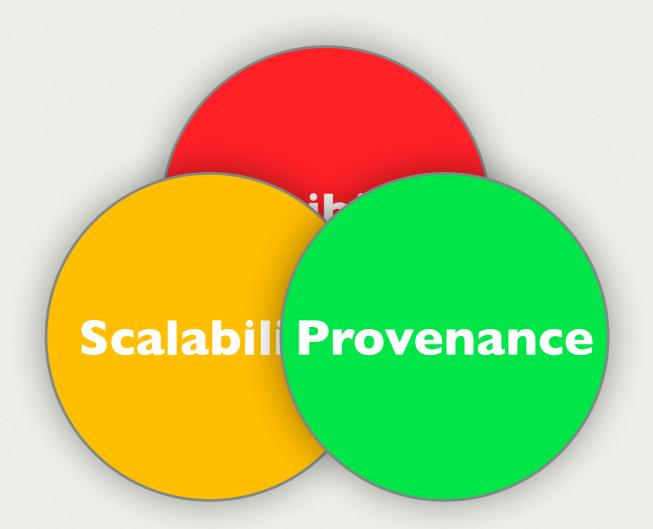




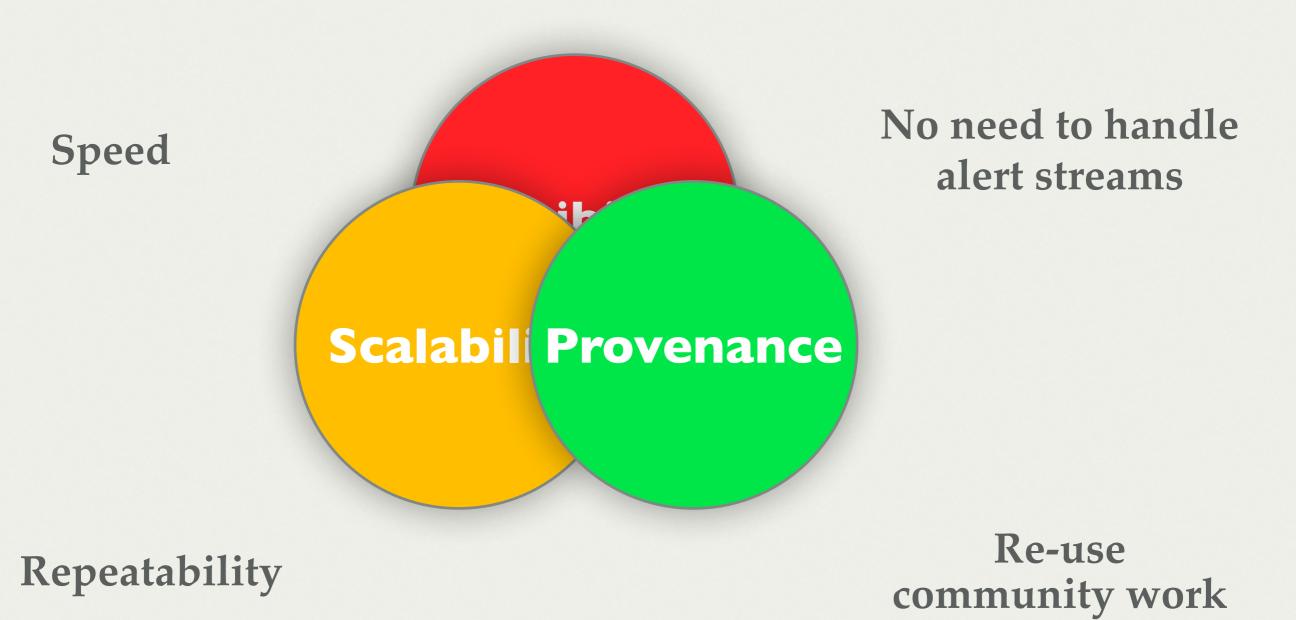




AMPEL



AMPEL



JOIN US!

GETTING IN

LEVEL: USER

- Define your own channel
 - Use existing units
 - Define run configurations

https://github.com/AmpelProject/Ampel-contrib-sample ampel-info@desy.de **GETTING IN**

LEVEL: *CONTRIBUTOR*

• Write (and share) T2 / T3 units

https://github.com/AmpelProject/Ampel-contrib-sample ampel-info@desy.de

GETTING IN

LEVEL: BUILDER

• Help us further developing core and instrument plugins

https://github.com/AmpelProject/Ampel-contrib-sample ampel-info@desy.de

THANKS

A&A in press https://arxiv.org/abs/1904.05922